

Coagulation of Dust - Collisions in the Decimetre Range

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ESSEN

Offen im Denken

- Planets form by accretion of km size protoplanetary bodies
- these planetesimals grow from dust particles
- different models of growth
 - coagulation of dust agglomerates
 - gravitational clumping of solids
 - combination of both
- decimetre size bodies important in both models
 - direct precursors of metre size bodies
 - can easily be trapped in vortices where gravitational collapse can occur

Introduction

- collision properties of decimetre size agglomerates important
- threshold conditions for fragmentation crucial for coagulation models

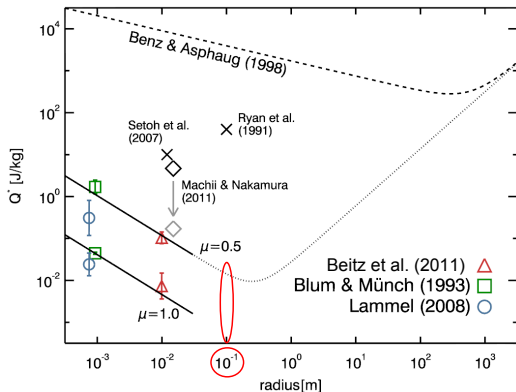


Figure : Critical fragmentation strength Q^* [Beitz et al. (2011)]

Experiment

Pressing of the Dust Agglomerates

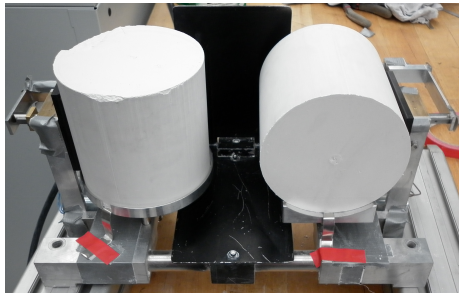
- used analogue material:
Quartz (SiO_2), irregular grains
- grain size: 0.1 to 10 μm , 80% of mass are in 1-5 μm (producer: Sigma-Aldrich)
- volume filling factor and grain size relevant for mechanical properties (Meisner et al. (2012), Schröppler et al. (2012), Blum et al. (2006))
- agglomerates with the same size ($\varnothing=12\text{cm}$) and volume filling factors ($\phi \approx 0.44$)
- mass of each agglomerate is $\approx 1.5\text{ kg}$



Experiment

Collision Setup

- experiment in a vacuum chamber in the drop capsule ($p \lesssim 10^{-2}$ mbar)
- carried out at drop tower in Bremen (4.7 s of microgravity)
- acceleration by a linear motor
- observation with two high speed cameras (500 fps)



Collision Videos

Example for Bouncing

- collision velocities between 0.8 and $7 \frac{\text{cm}}{\text{s}}$

Collision Videos

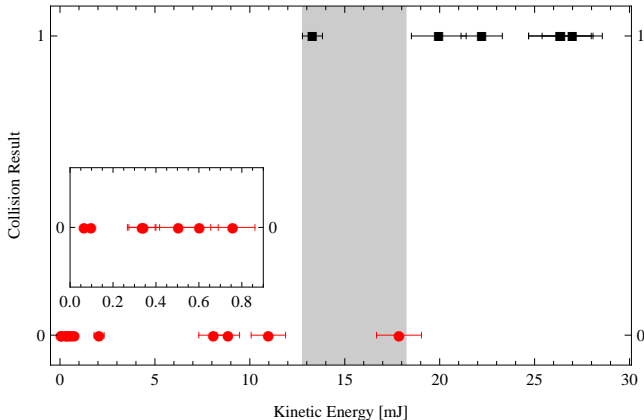
Example for Fragmentation

- collision velocity at $\approx 25.7 \frac{\text{cm}}{\text{s}}$

First Results

Critical Collision Velocities

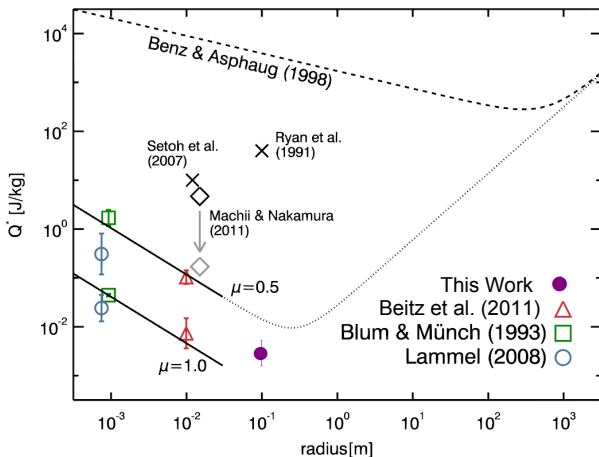
- Bouncing = 0, Fragmentation = 1
- critical fragmentation velocity at about $16 \frac{\text{cm}}{\text{s}}$



First Results

Critical Fragmentation Strength

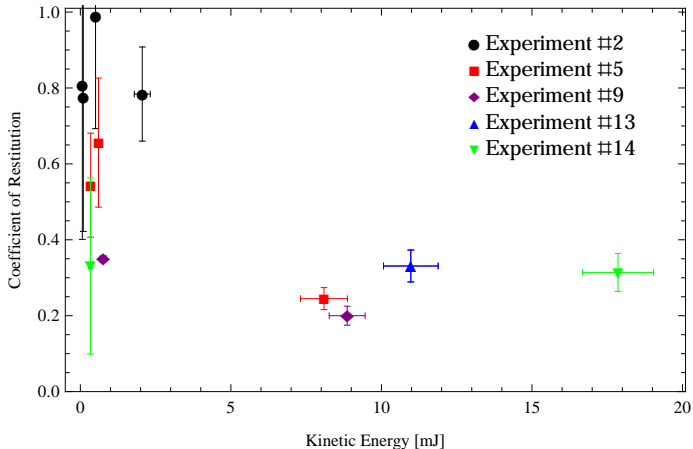
- $\mu=1$: boundary between bouncing and fragmentation
- critical fragmentation strength $Q^* \approx (5 \pm 0.6) \cdot 10^{-3} \frac{\text{J}}{\text{kg}}$



First Results

Coefficient of Restitution

- ratio of kinetic energy after and before the collision
- decreases with increasing collision energy
- dependent on primary collision energy



Conclusions

- critical fragmentation velocity at $\approx 16 \frac{\text{cm}}{\text{s}}$ ($Q^* \approx (5 \pm 0.6) \cdot 10^{-3} \frac{\text{J}}{\text{kg}}$)
- elastic behaviour at low collision velocities
- coefficient of restitution decreases with increasing collision energy