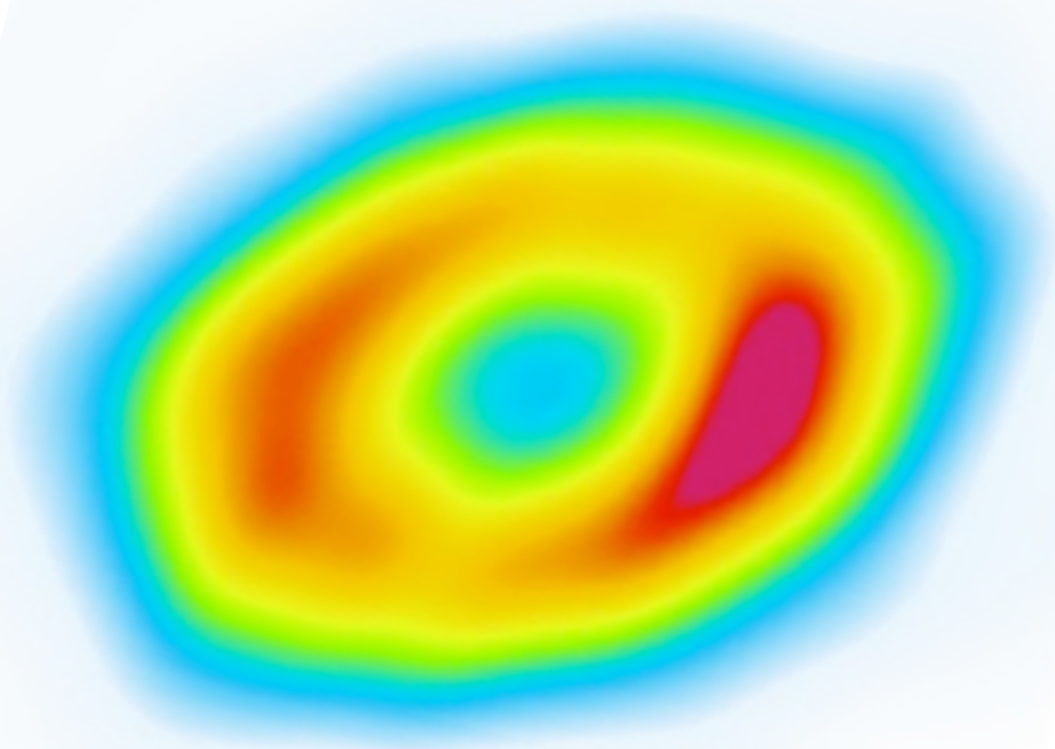


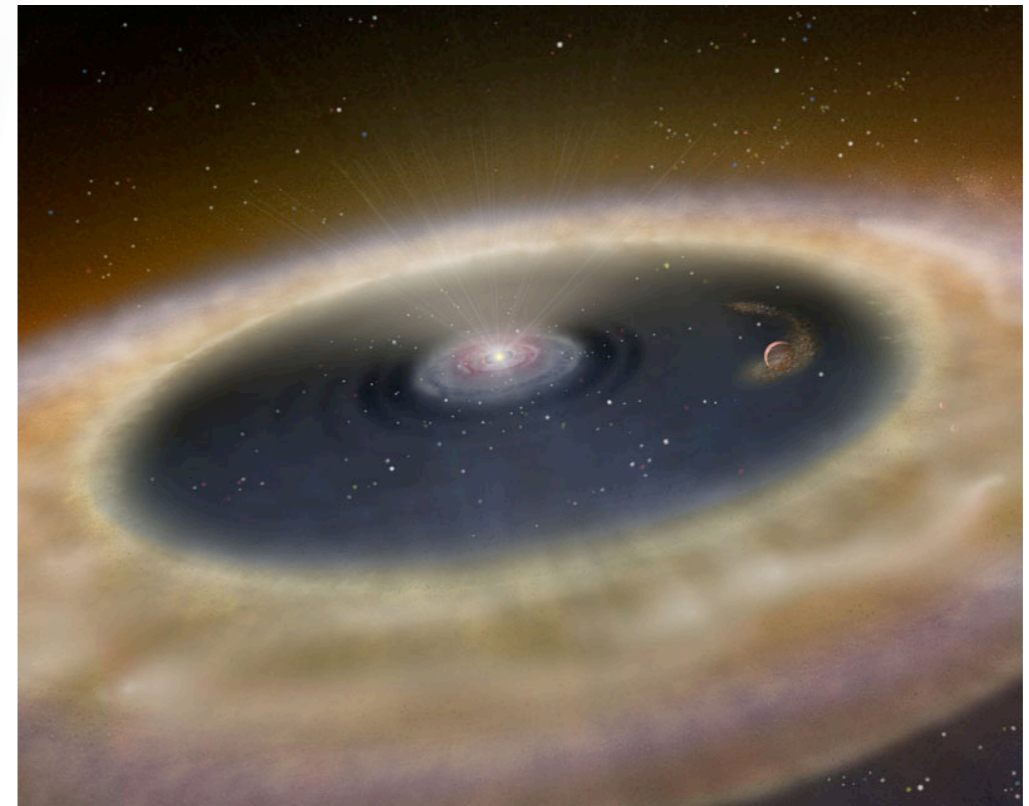
A Millimeter-Wave View of the “Transition” Disks

Sean Andrews

Harvard-Smithsonian Center for Astrophysics

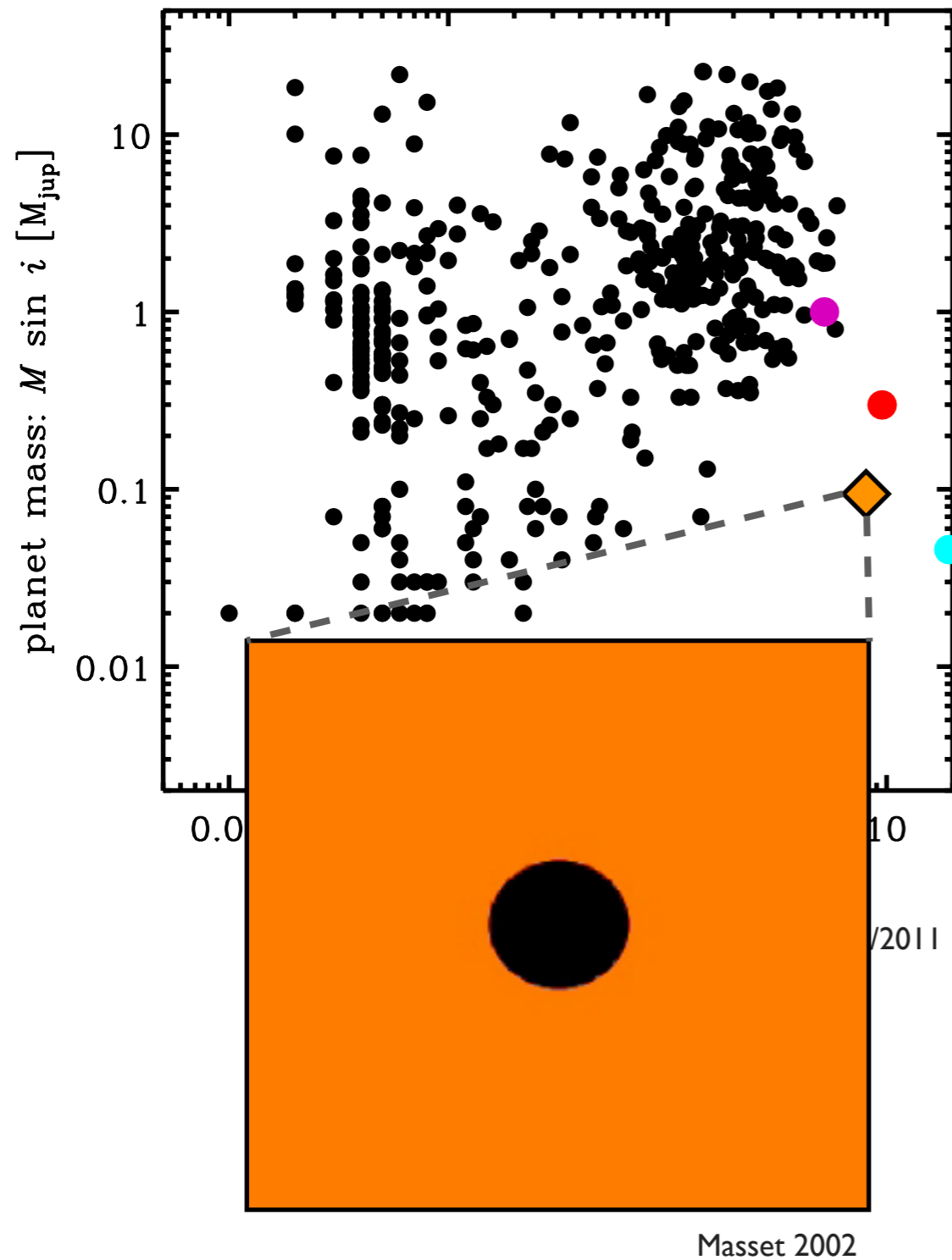


Andrews et al 2011b; SMA + PdBI (0.88mm)



Kraus et al 2012; illustration by K Teramura

planetary origins: disks as “initial conditions”



1. [feasibility](#) (disk properties)

Q: is there enough stuff(r,t) ?

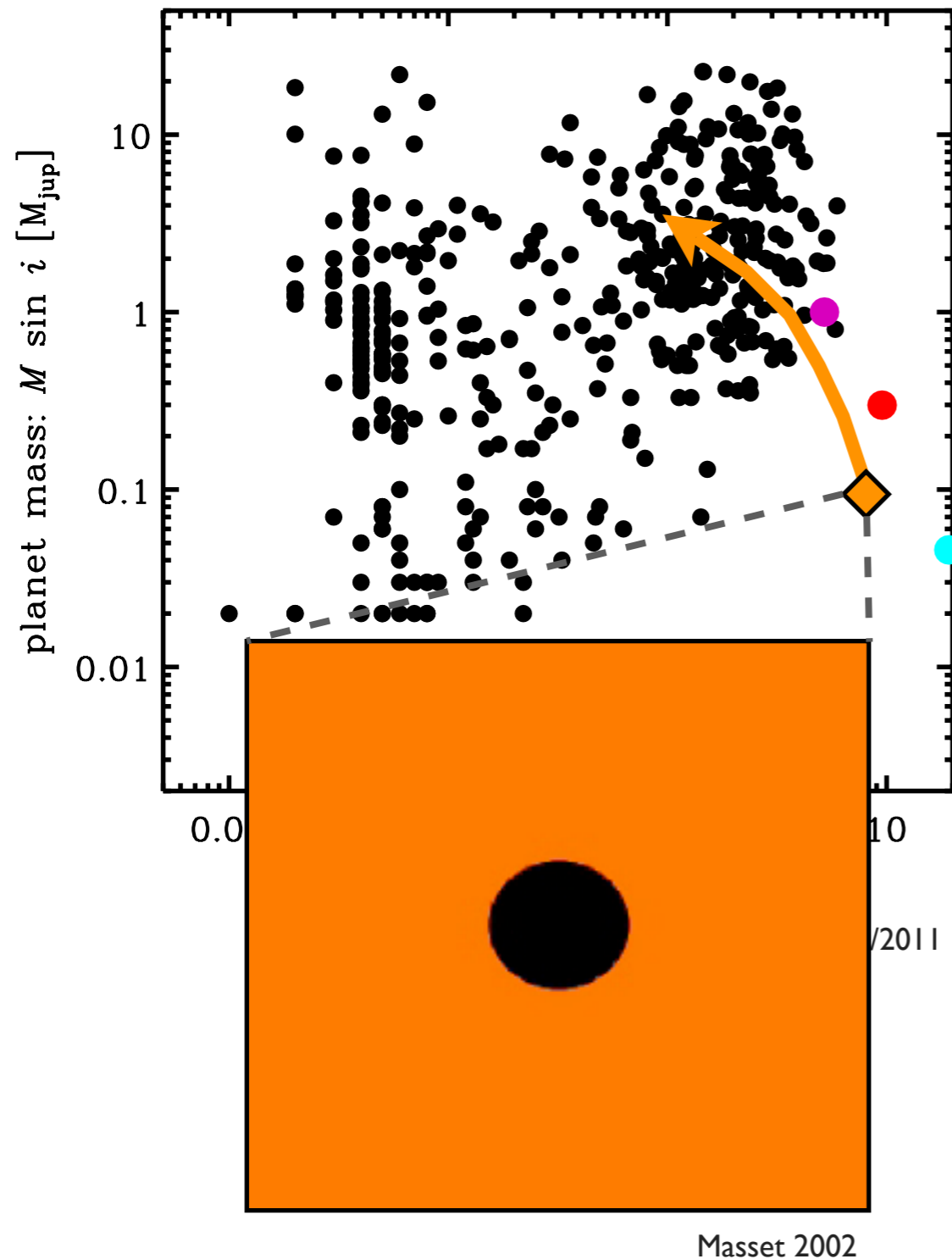
A: yes, probably

2. [early evolution](#) (disk+planets)

Q: how does planet($r,t,disk$) ?
or disk($r,t,planets$) ?

A: ???...

planetary origins: disks as “initial conditions”



1. [feasibility](#) (disk properties)

Q: is there enough stuff(r,t) ?

A: yes, probably

2. [early evolution](#) (disk+planets)

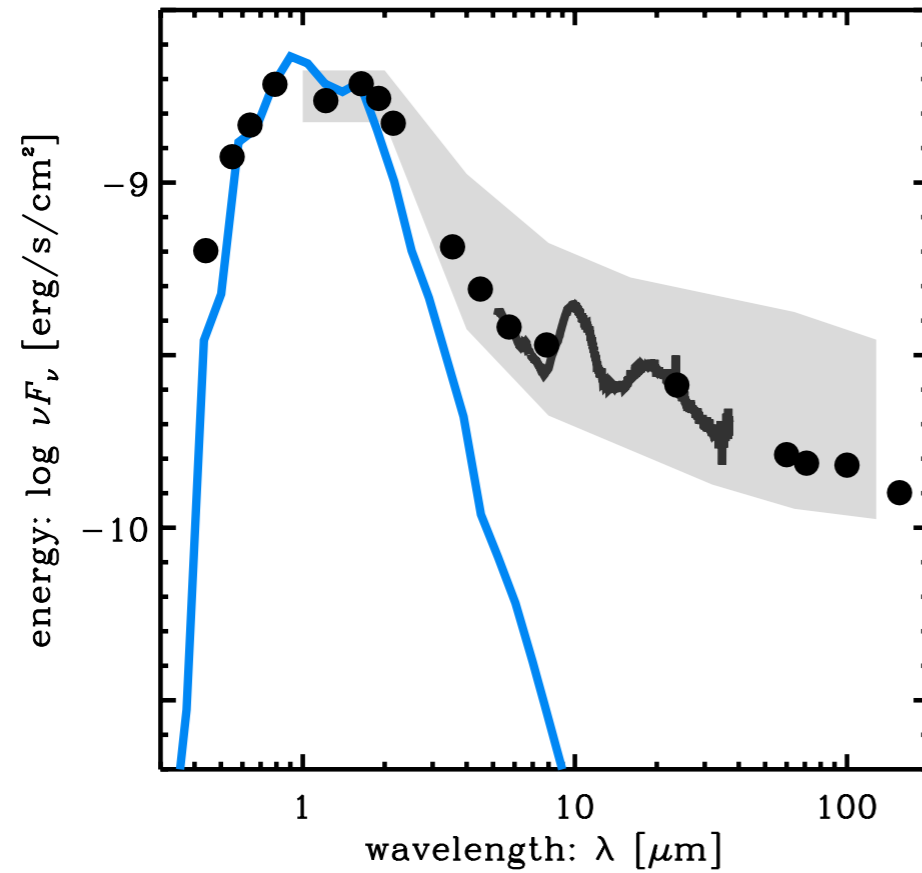
Q: how does planet($r,t,disk$) ?
or disk($r,t,planets$) ?

A: ???...

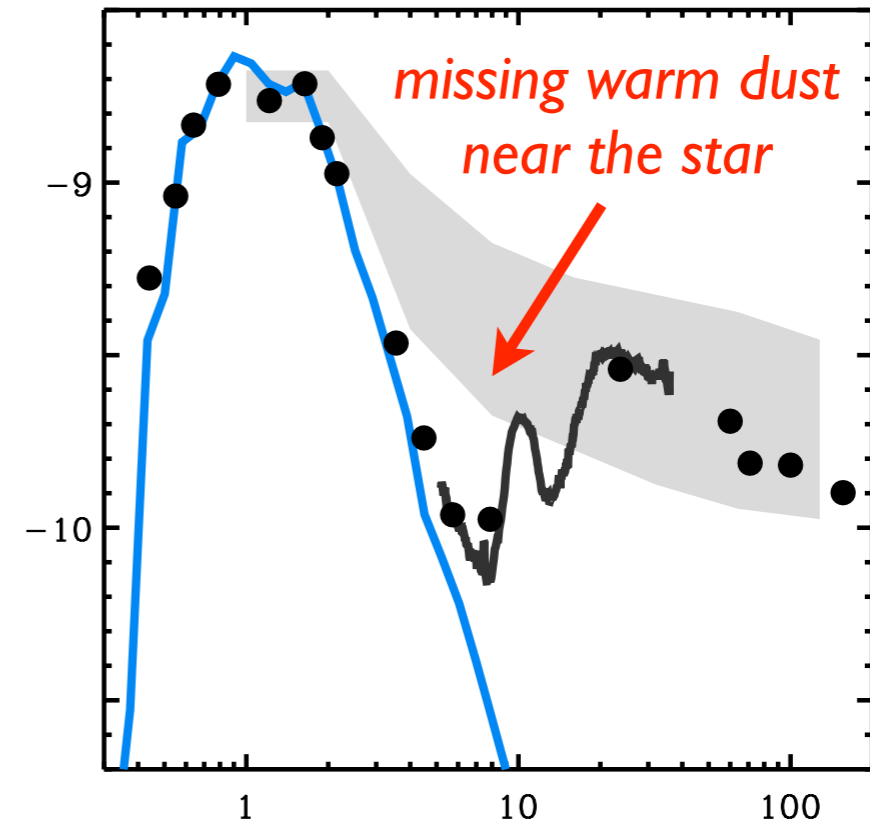
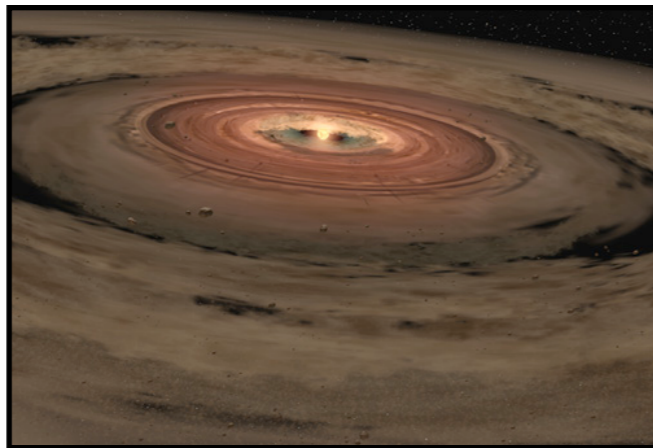
measured planet properties
are *not* their initial values

...[disk evolution](#) matters...

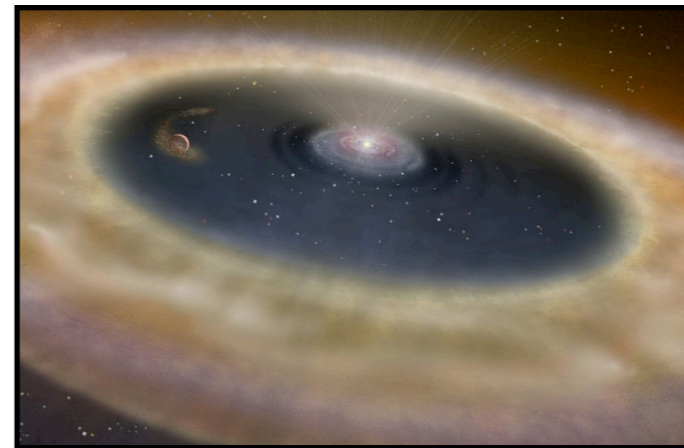
what is a transition disk?



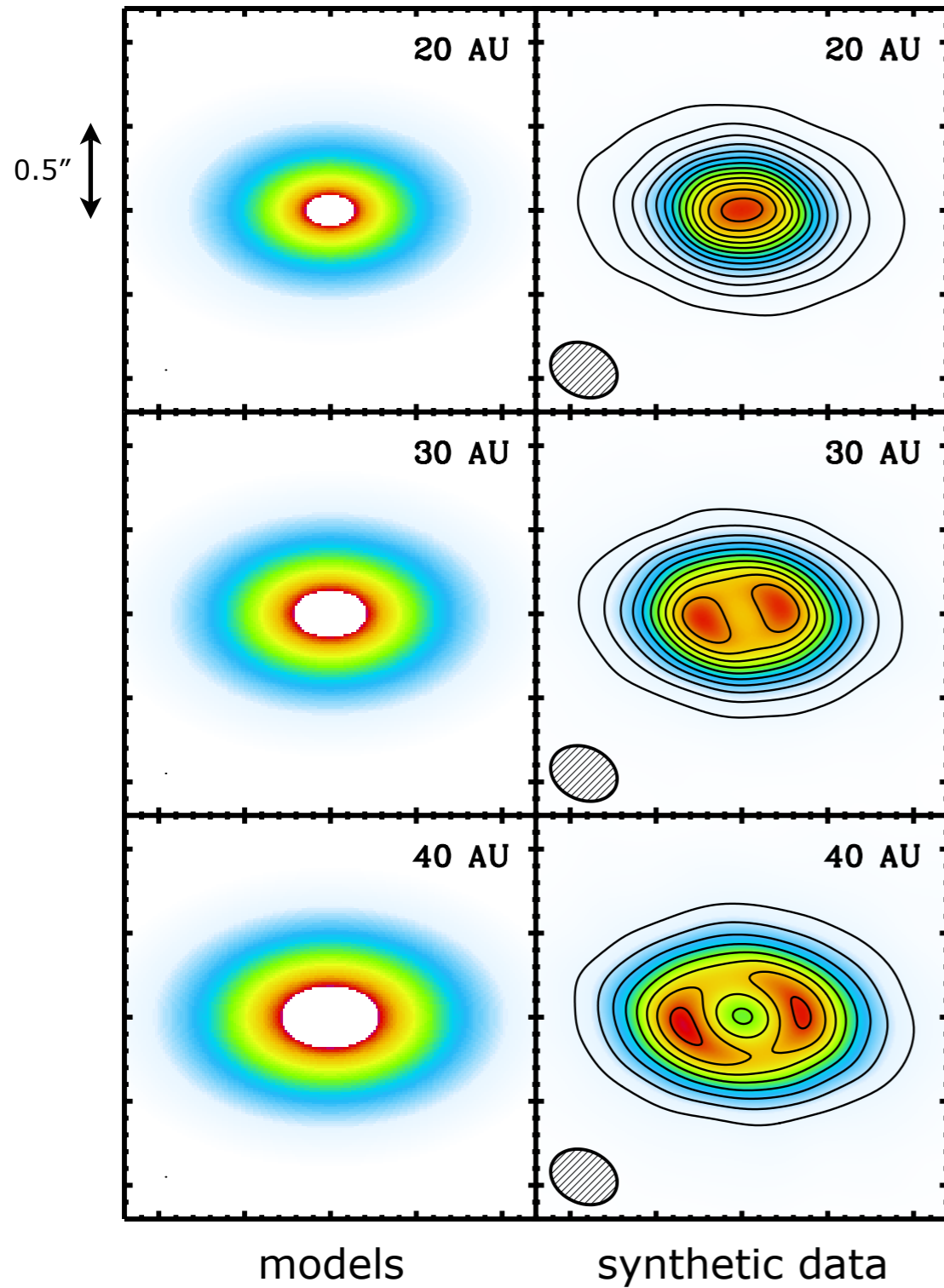
"normal" disk



"transition" disk

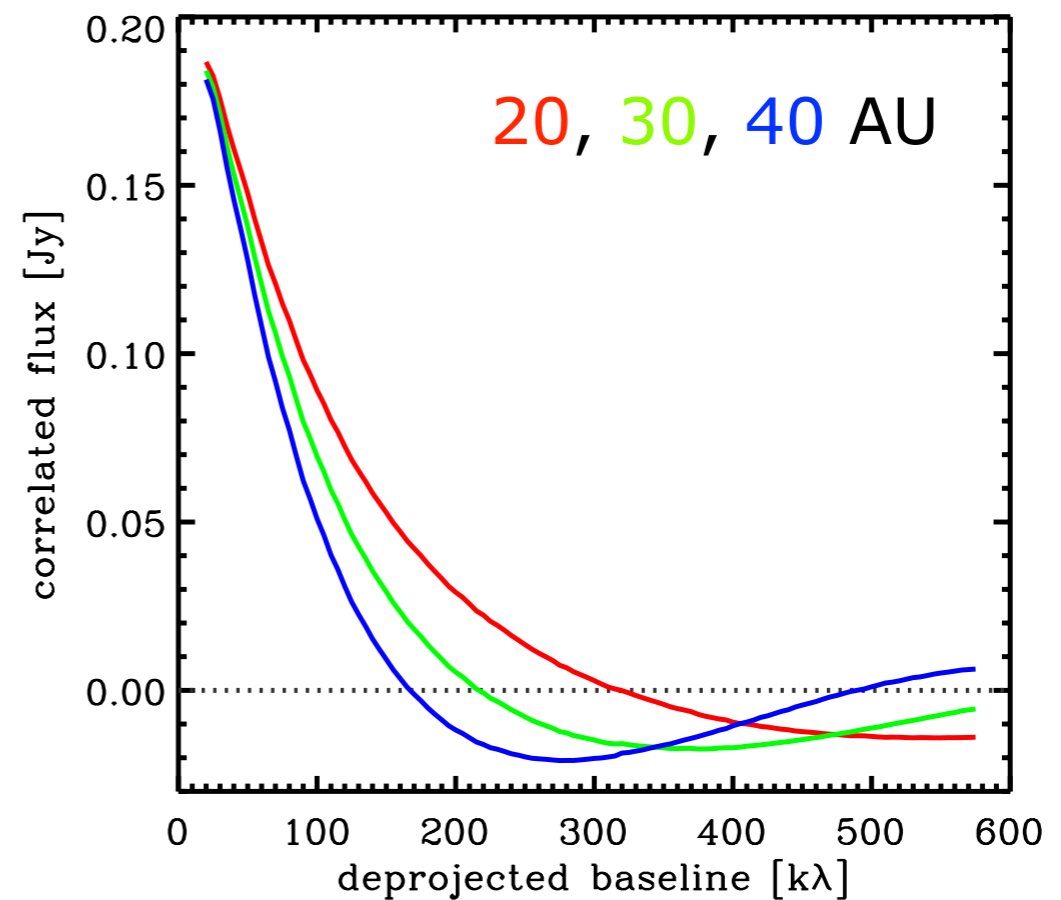


mm-wave imaging of transition disk "cavities"

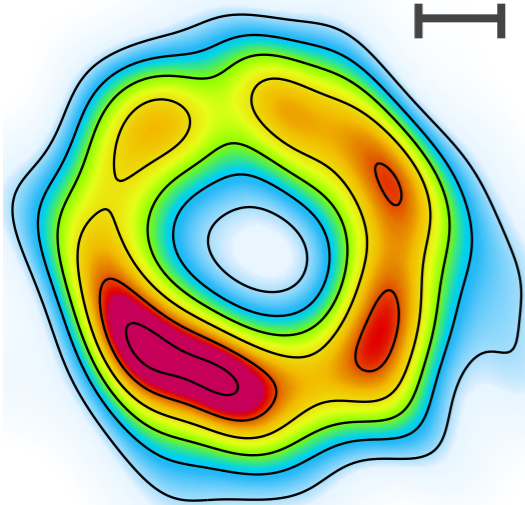


observing requirements:

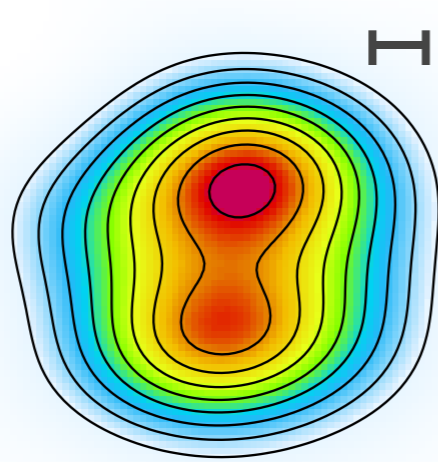
- high angular resolution ($< 0.5''$)
- good sensitivity (> 100 mJy)
- (helps to know viewing geometry)



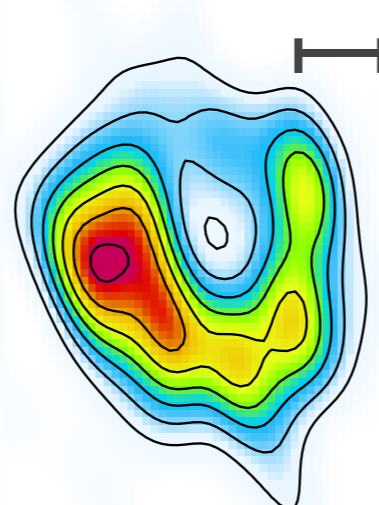
mm-wave transition disk images



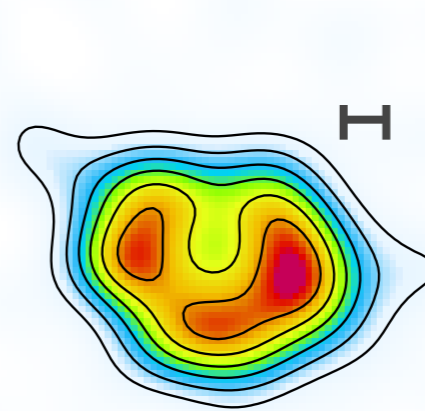
Mathews et al 2012



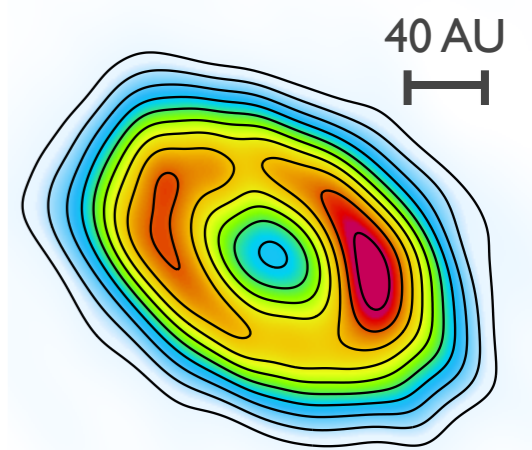
Isella et al 2010b



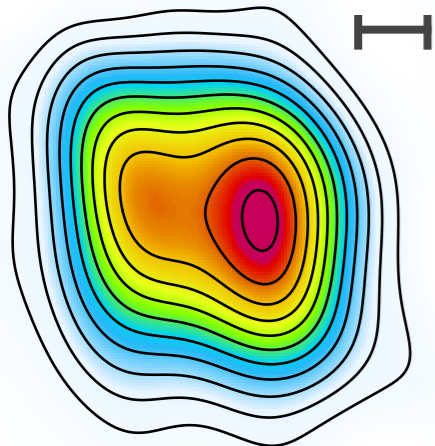
Brown et al 2008



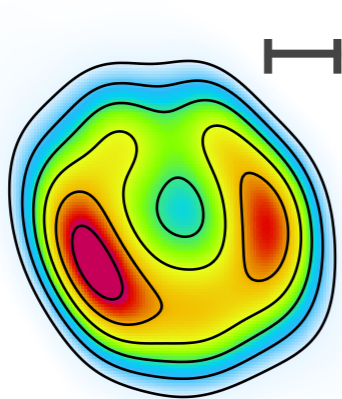
Brown et al 2008



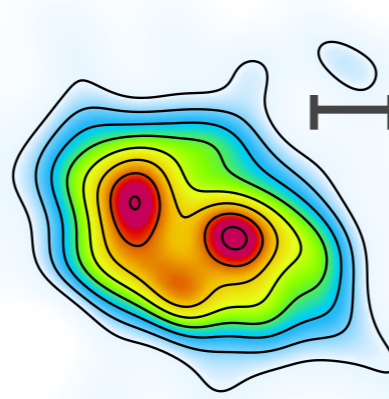
Andrews et al 2011b



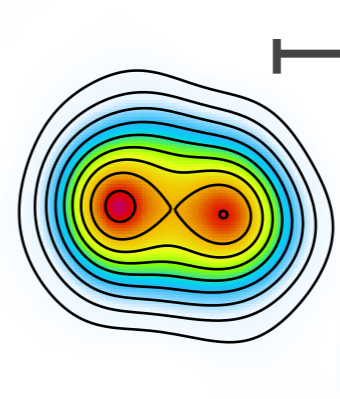
Andrews et al 2009



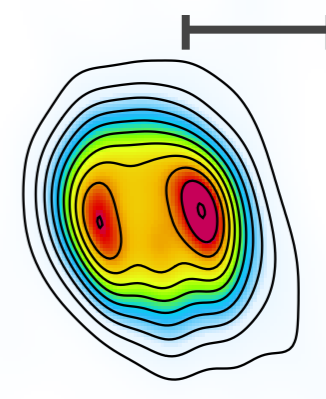
Andrews et al 2009



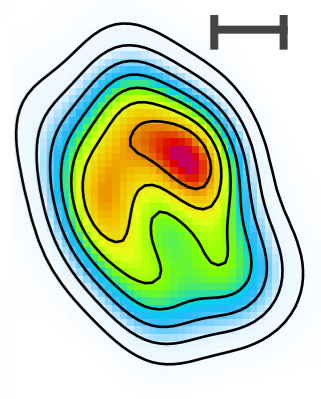
Hughes et al 2009



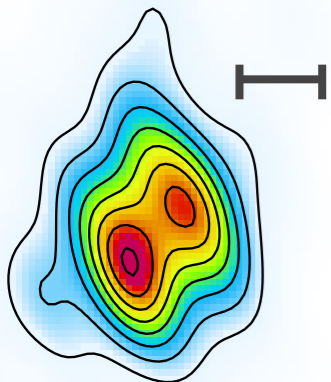
Cieza et al 2012



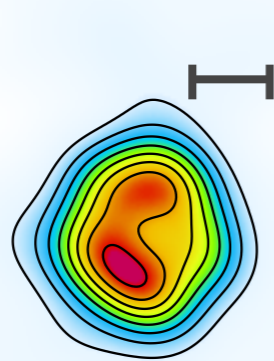
Rosenfeld et al 2012



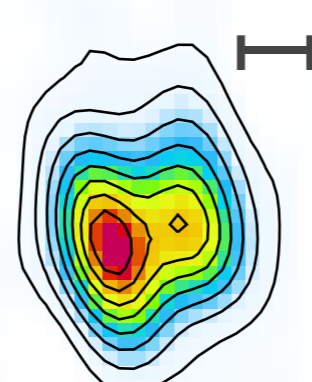
Andrews et al 2010



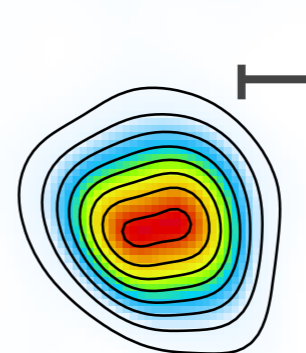
Andrews et al 2011a



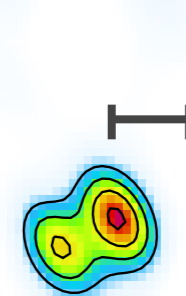
Andrews et al 2011a



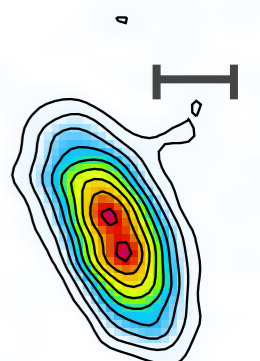
Andrews et al 2009



Brown et al 2012

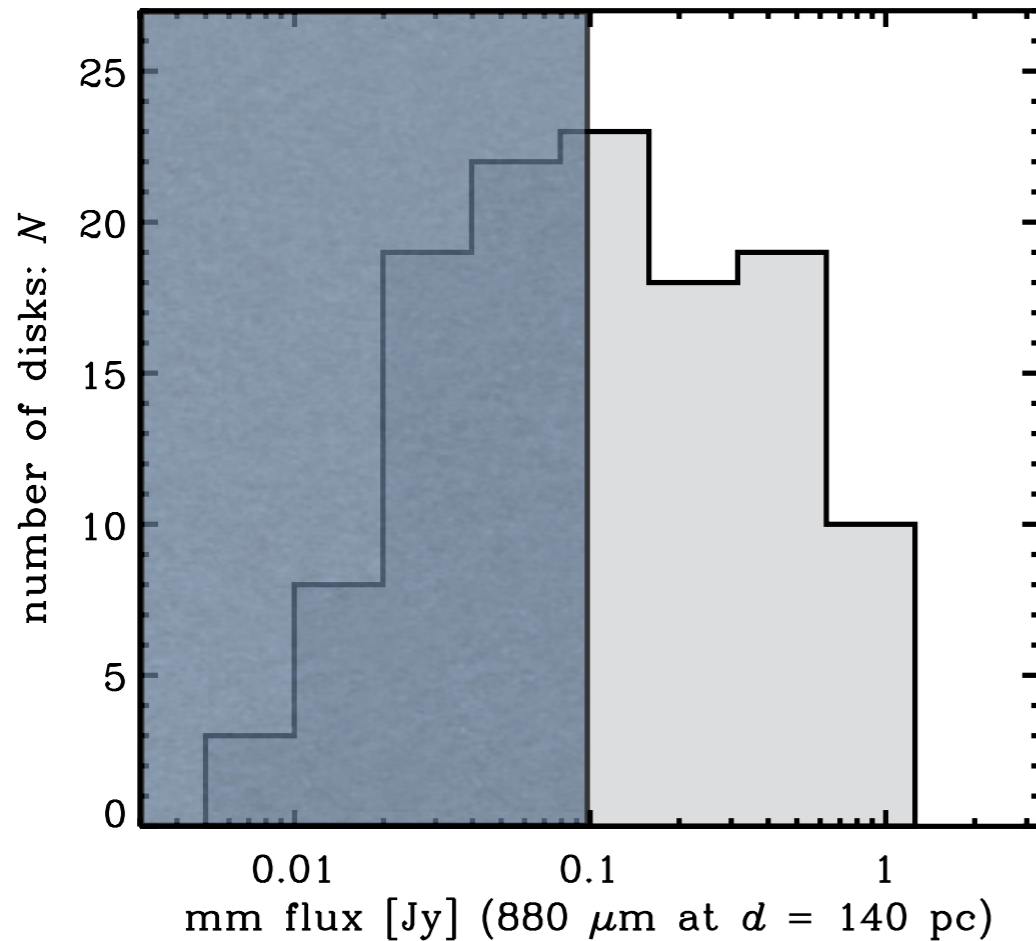


Andrews et al 2011a

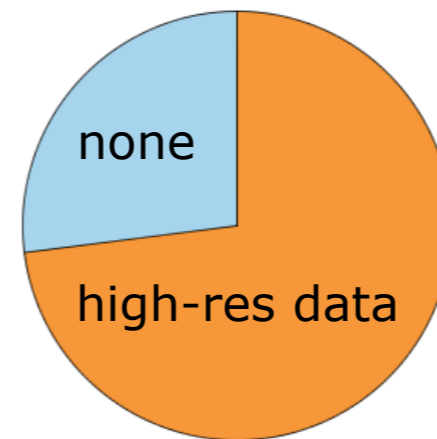


Isella et al 2010a

how common is the transition disk phenomenon?

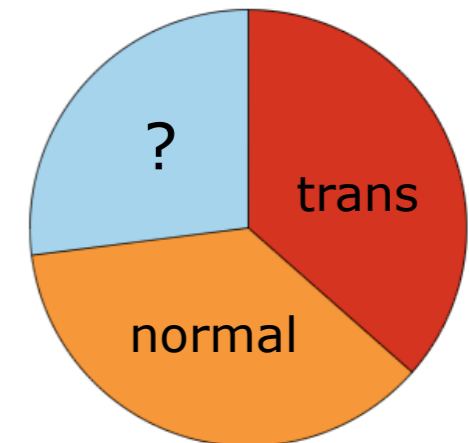


Andrews & Williams 2005, 2007; Nurnberger et al 1998



*3/4 of bright disks
have been imaged
at $<0.5''$ resolution*

Andrews et al 2009-2012
Isella et al 2009-2012
Guilloteau et al 2012



*>1/3 of bright disks
have large cavities
($R > 15$ AU)*

Andrews+ 2011; Brown+ 2008, 2012
Isella+ 2010, 2012; Cieza+ 2012
Mathews+ 2012; Rosenfeld+ 2012

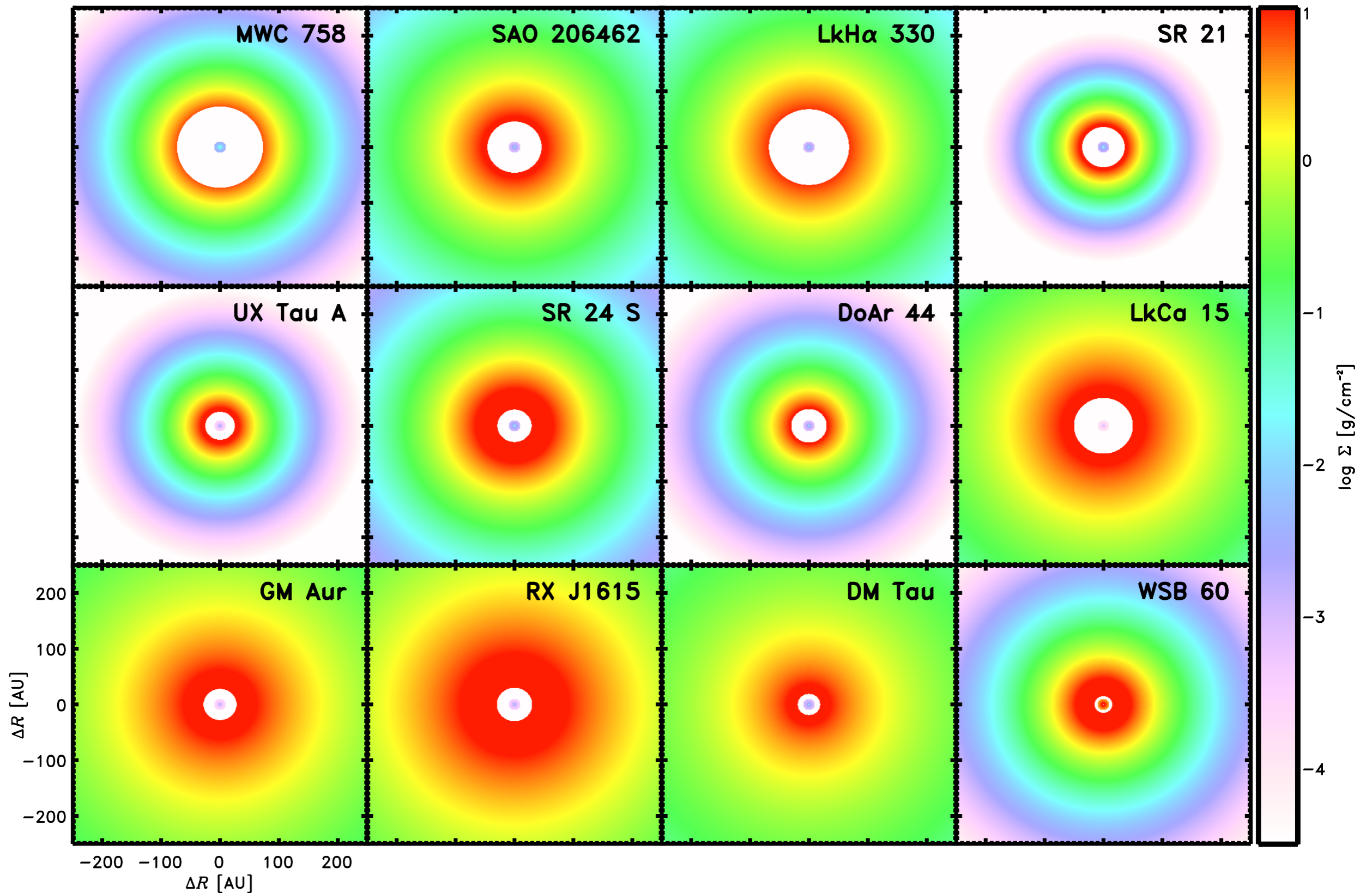
what fraction have cavities?

a disk mass threshold?

preferential stellar properties?

correlations with cavity size?

modeling transition disk structures



basic transition disk properties

1. transition disks are *common*

- > 1/3 of bright (massive) disks
- bias: large, tenuous cavities

2. observed cavities are *large*

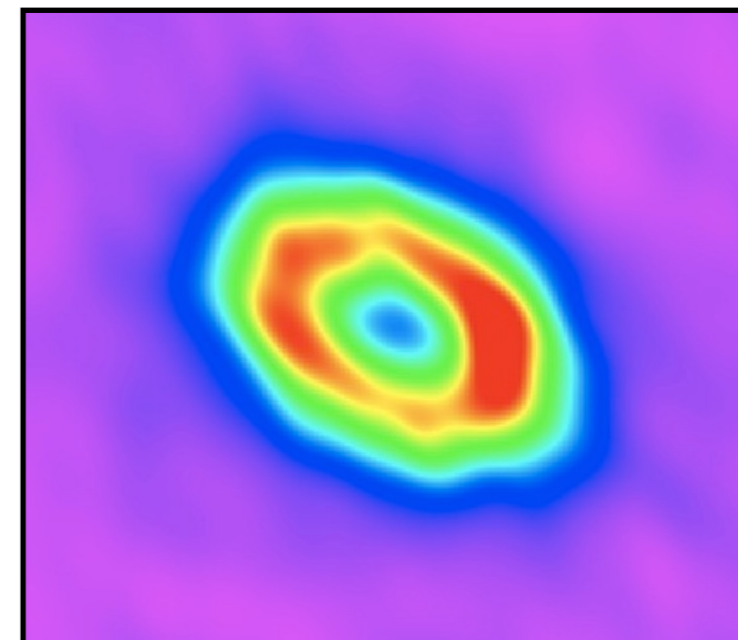
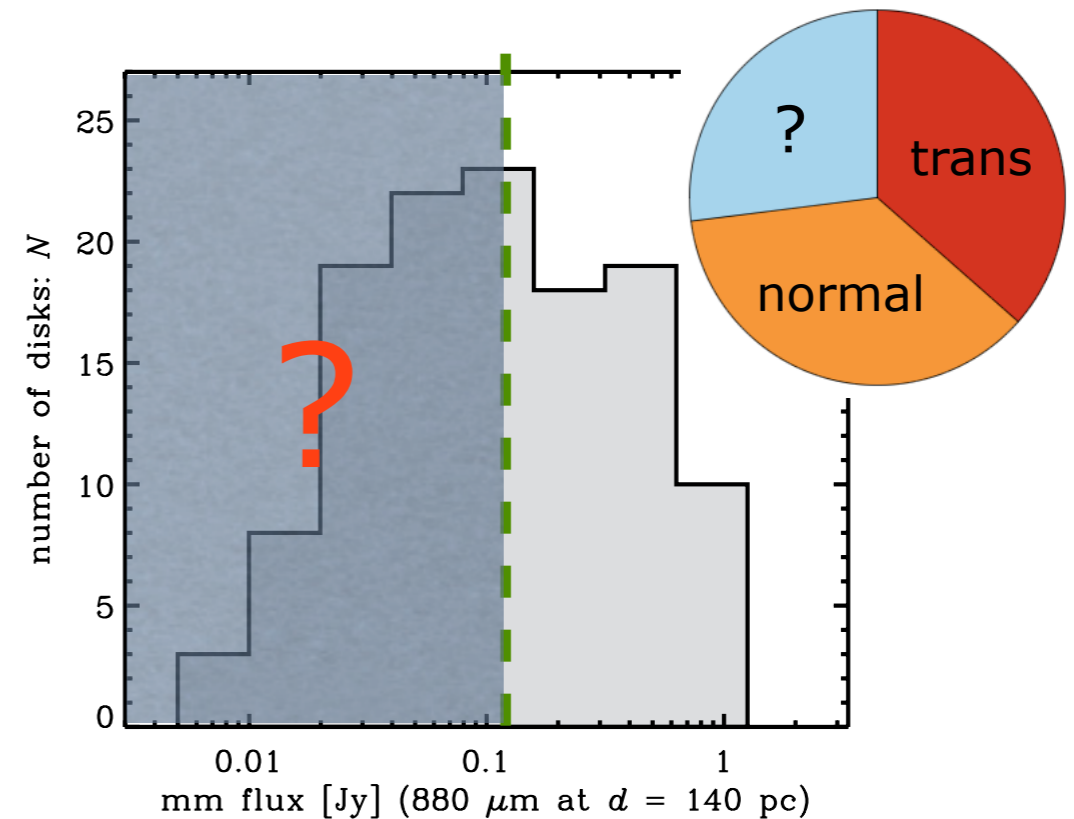
- radii of 15-75 AU (or more)

3. cavities are *depleted*

- $\tau(\text{IR})$ down by $10^{3-5} \times$
- $\tau(\text{mm})$ down by $> 100 \times$
- but they are *not* empty

4. outer regions like normal disks

- typical density profiles, sizes



low-mass (planetary?) companions

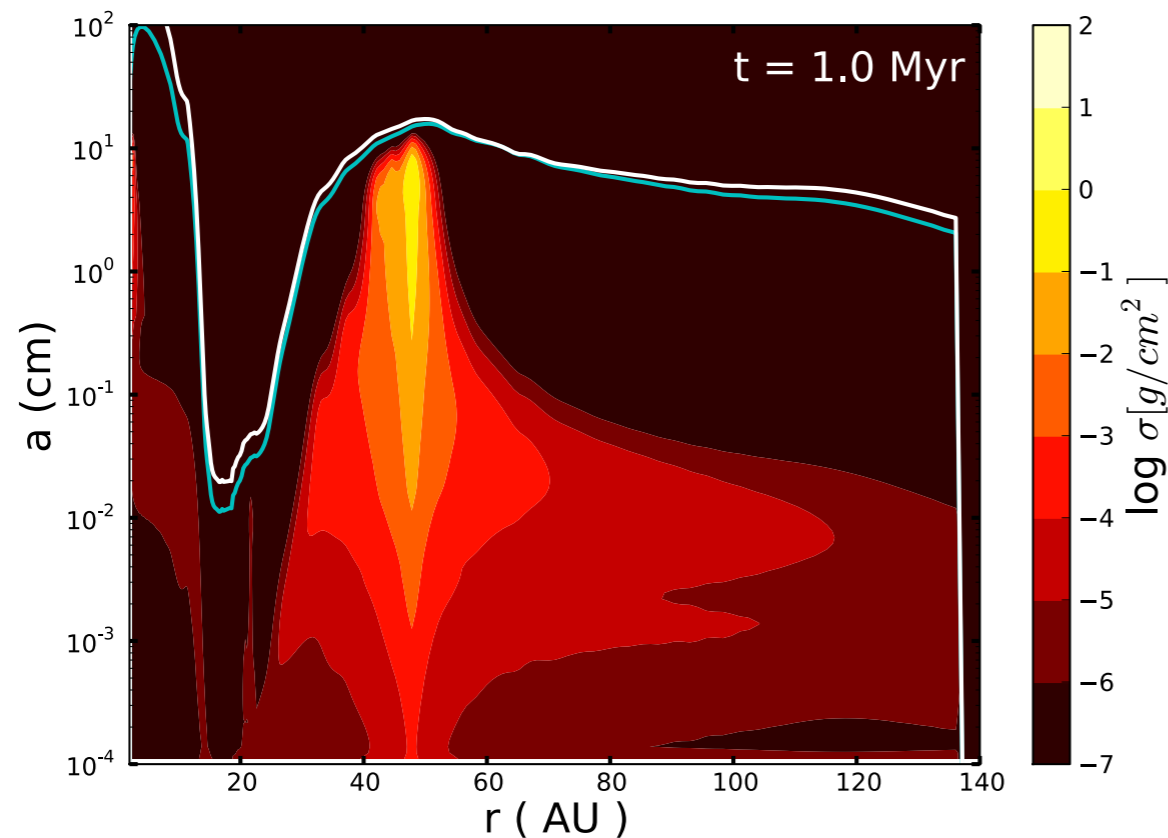
resonant torques make waves

gap opens; pressure bump

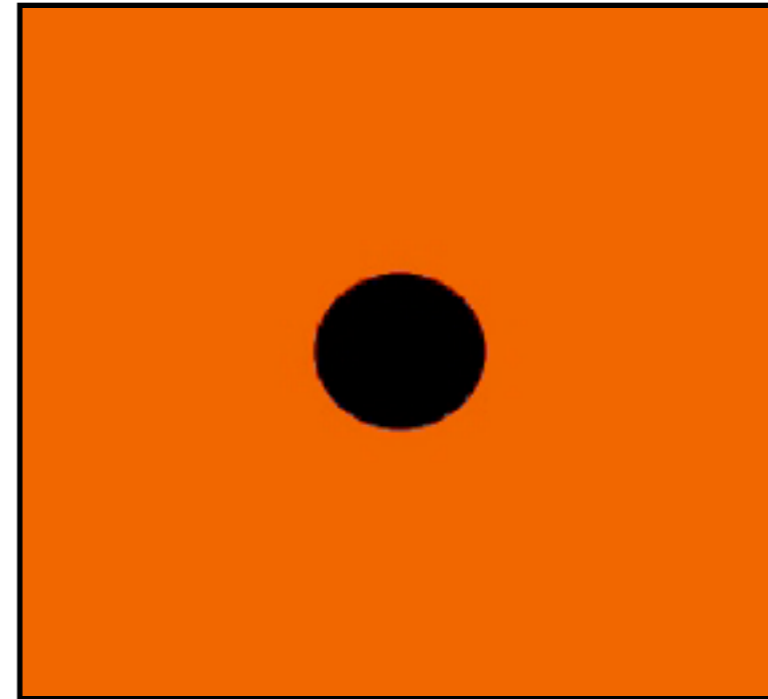
Lin & Papaloizou 1979; Bryden et al 1999

key issue: *dust/gas flow across gap*

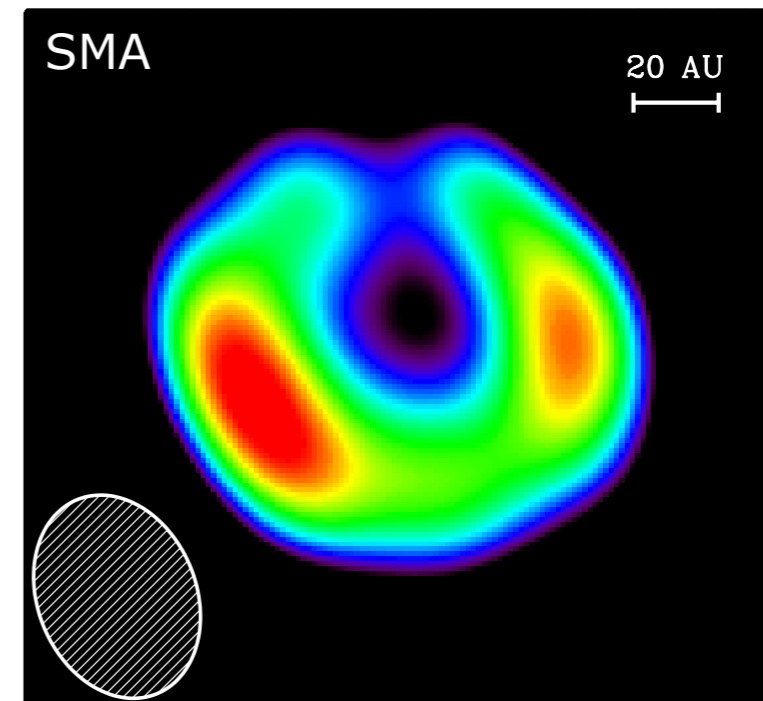
we ought to reconsider modeling formalism



Pinilla, Benisty, & Birnstiel 2012

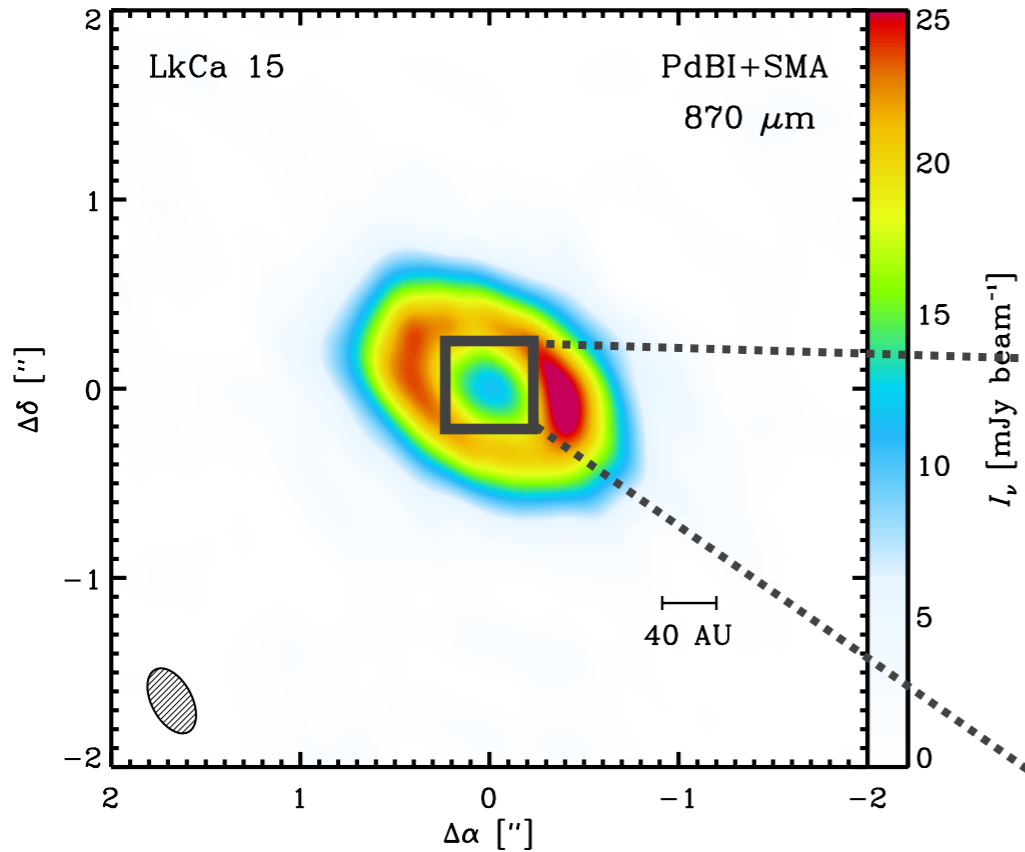


Masset 2002

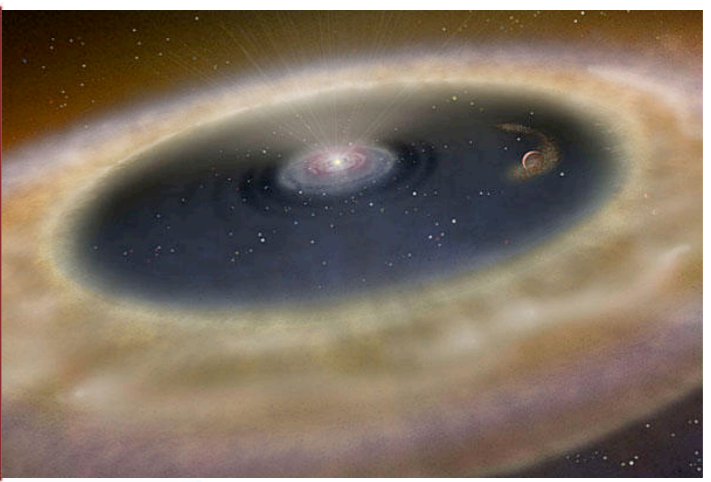
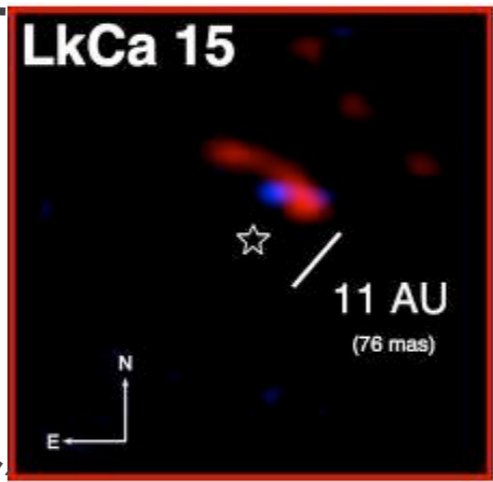


Andrews et al 2009

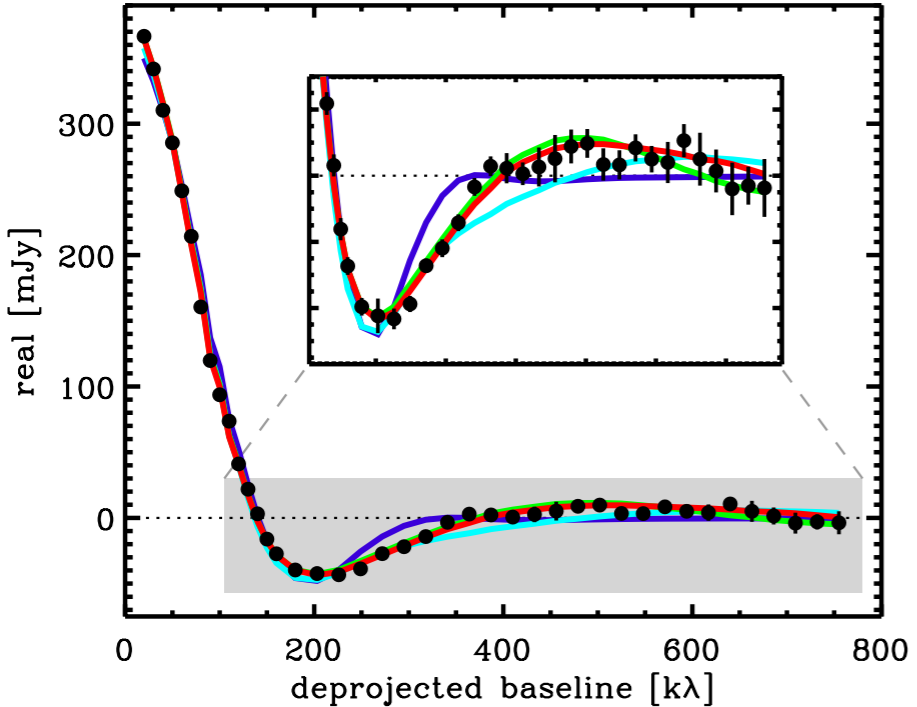
a (large) population of very young exoplanets?



semi-analytic disk gap model:
 4-9 M_{Jup} planet orbiting at ~ 16 AU
 Andrews, [Rosenfeld](#), et al 2011



Kraus & Ireland 2012



Lagrange et al 2009

Wilner, Andrews, & Hughes 2011

are these the progenitors of debris disks?

SUMMARY:

transition disks host young planetary systems (?)

1) transition disks have low optical-depth dust cavities:

imaging is critical; best at mm

2) a “transition” phase:

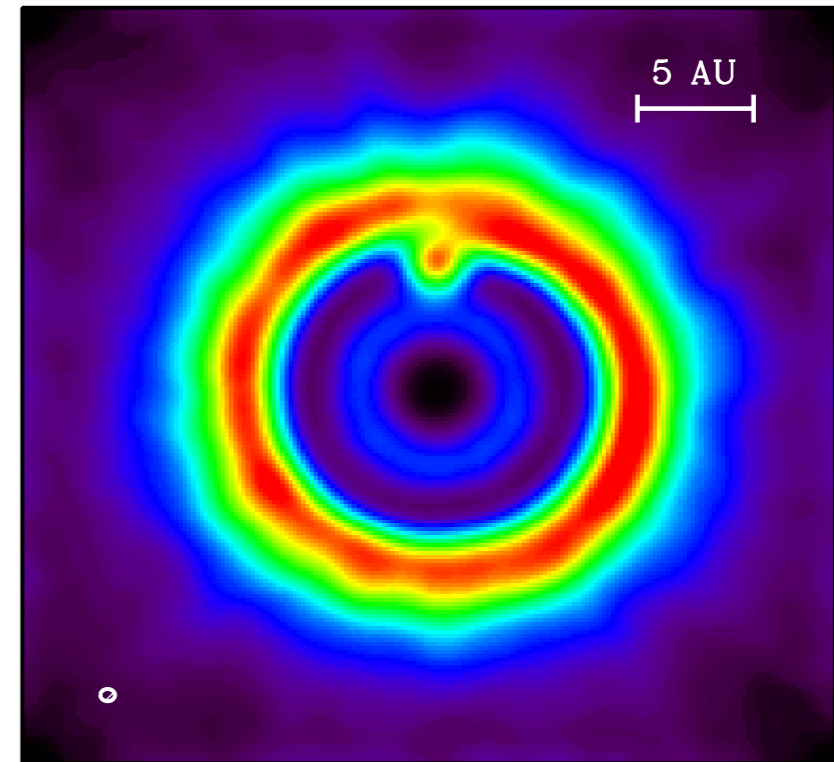
common, at least for massive disks

3) likely due to tidal interactions with faint companion(s)

mm/radio imaging can (indirectly) find/characterize young planets

cavity size \sim planet orbit

depletion + $\dot{M} \sim$ planet mass



bright future with ALMA

many thanks to:

David Wilner
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Andrea Isella
Lucas Cieza
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