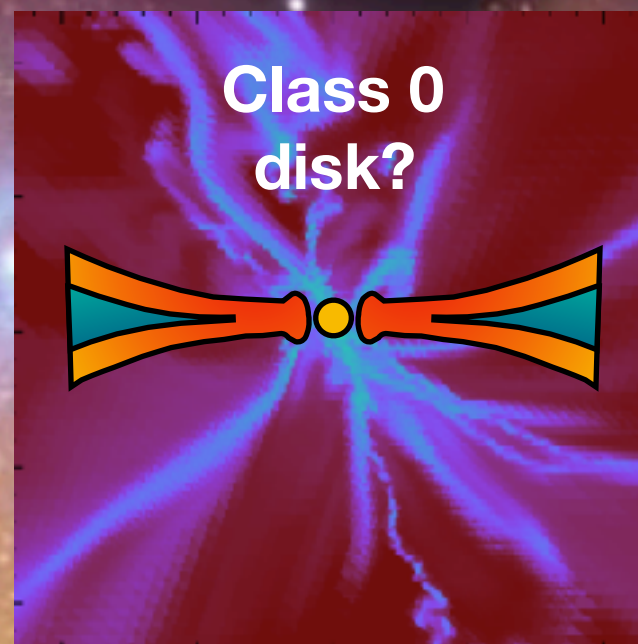




ALMA view of a Class 0 Protostellar system at disk scales

María José Maureira

Postdoc (CAS@MPE)



Class 0
Envelope

The laboratory: the deeply embedded Class 0 IRAS 16293-2422

0.14 pc

Herschel column density
Ladjelate et al. 2020

ALMA continuum

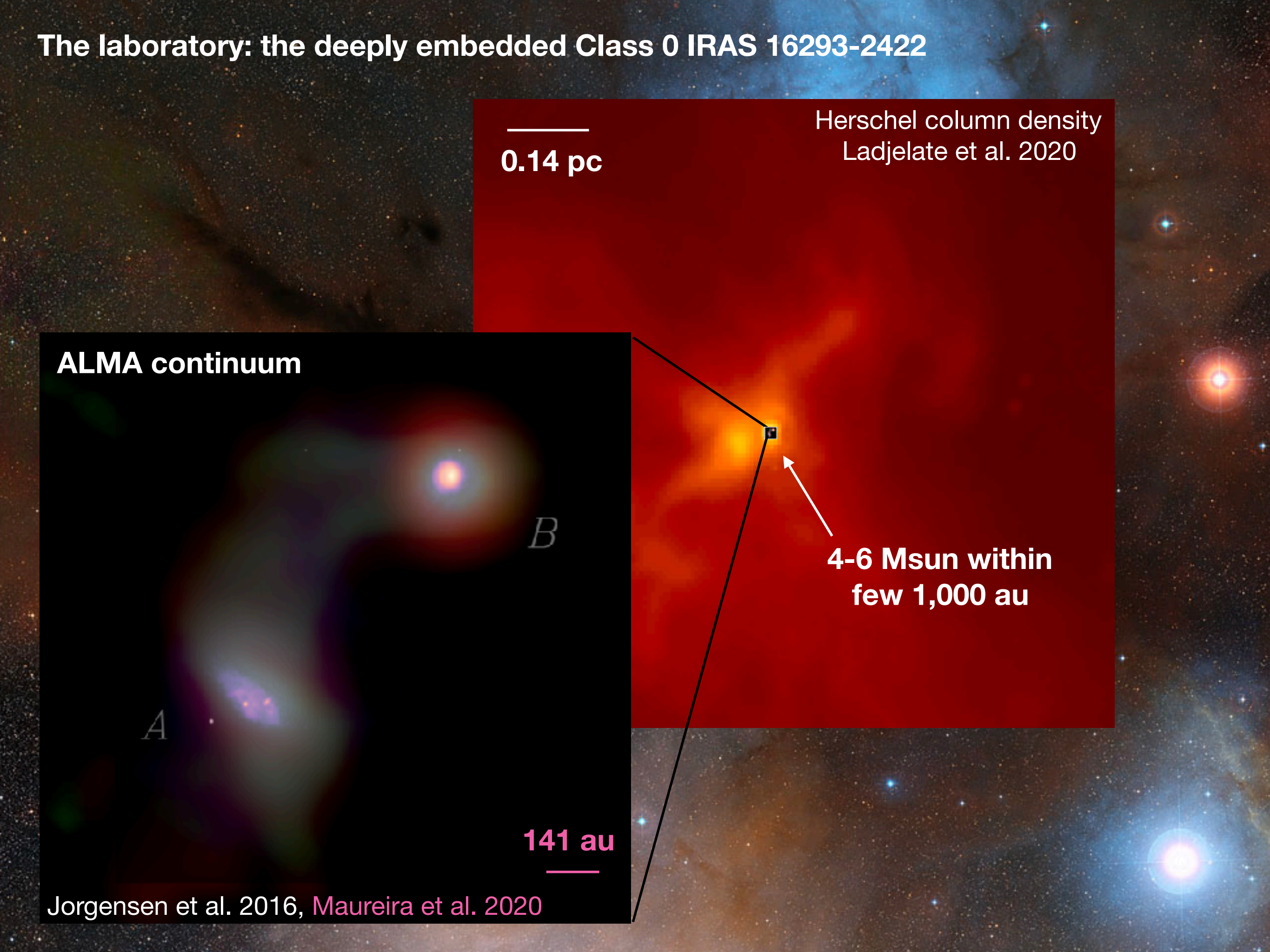
A

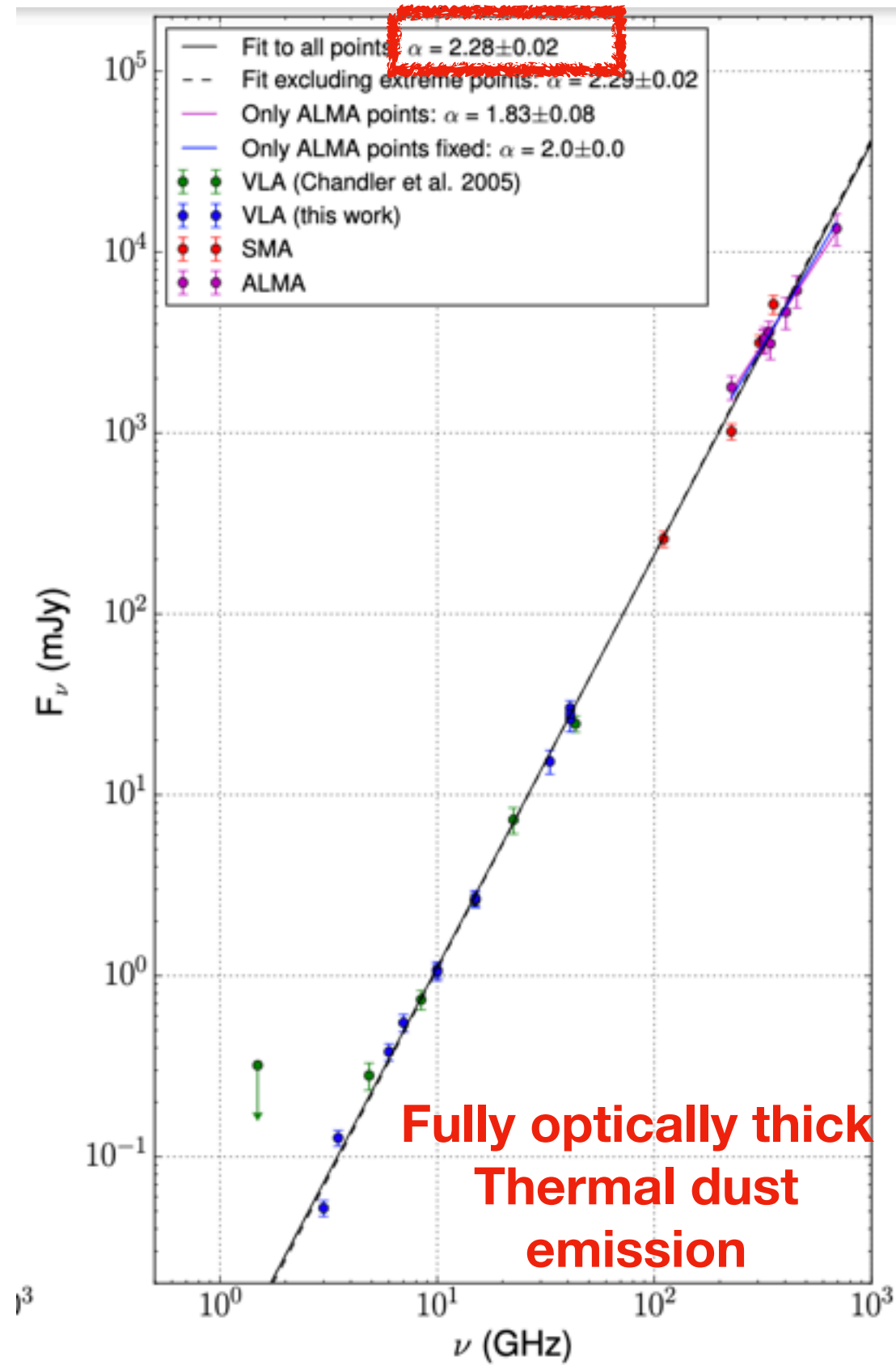
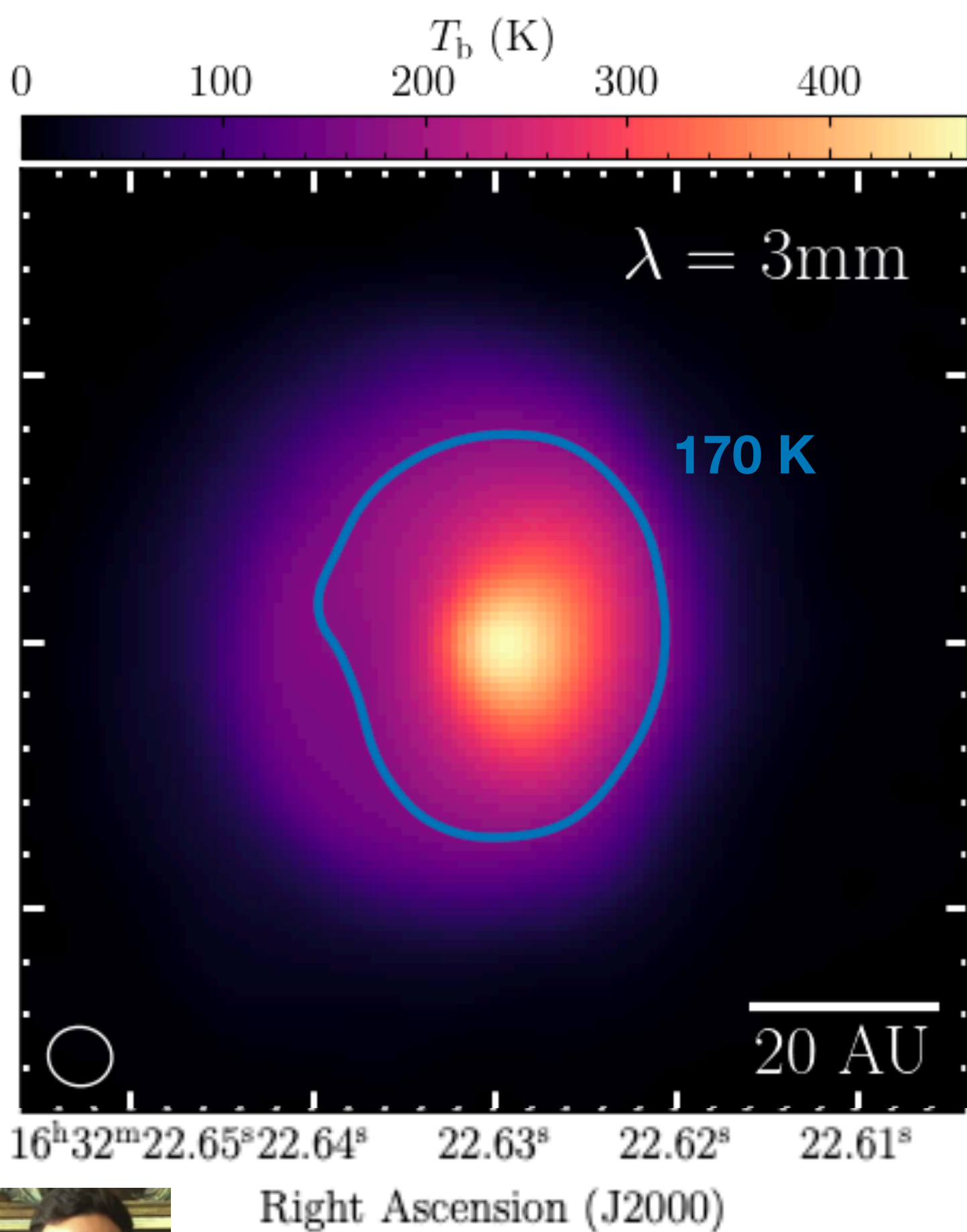
B

4-6 Msun within
few 1,000 au

141 au

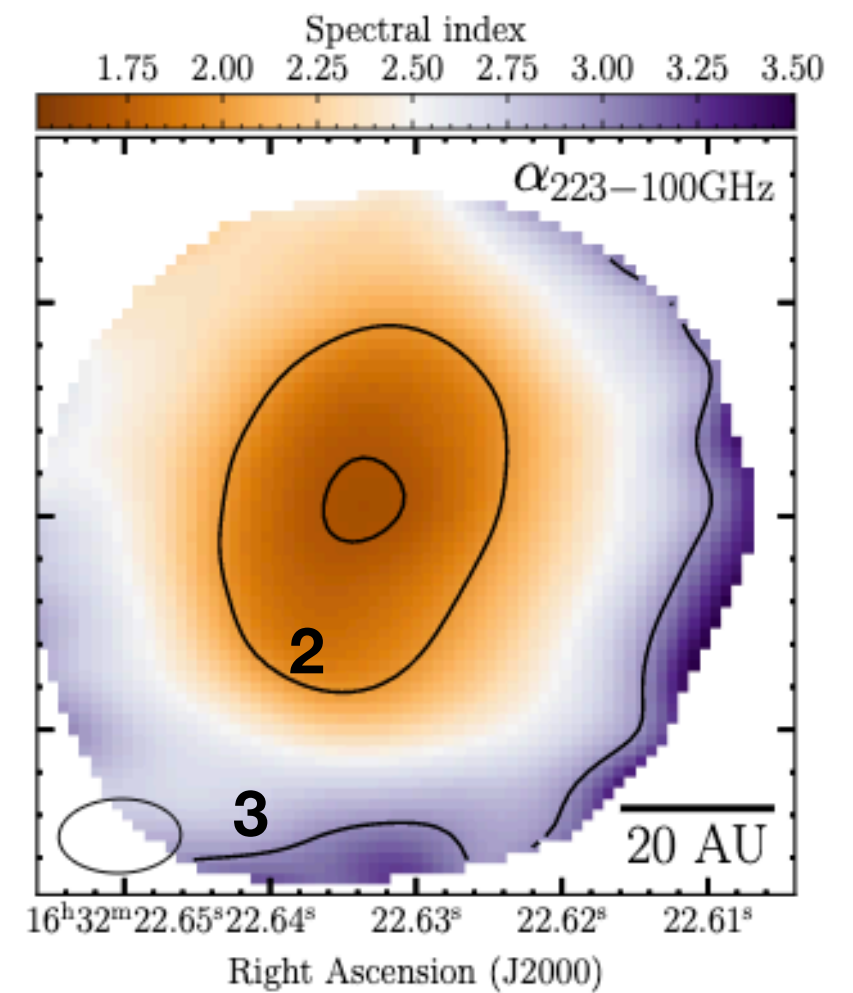
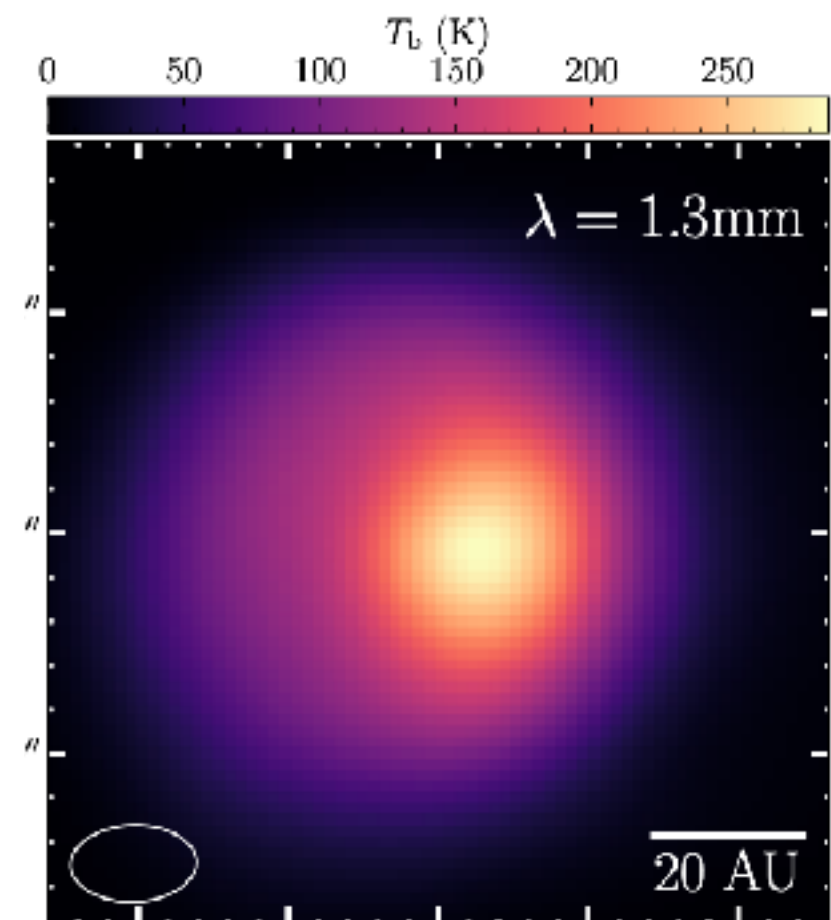
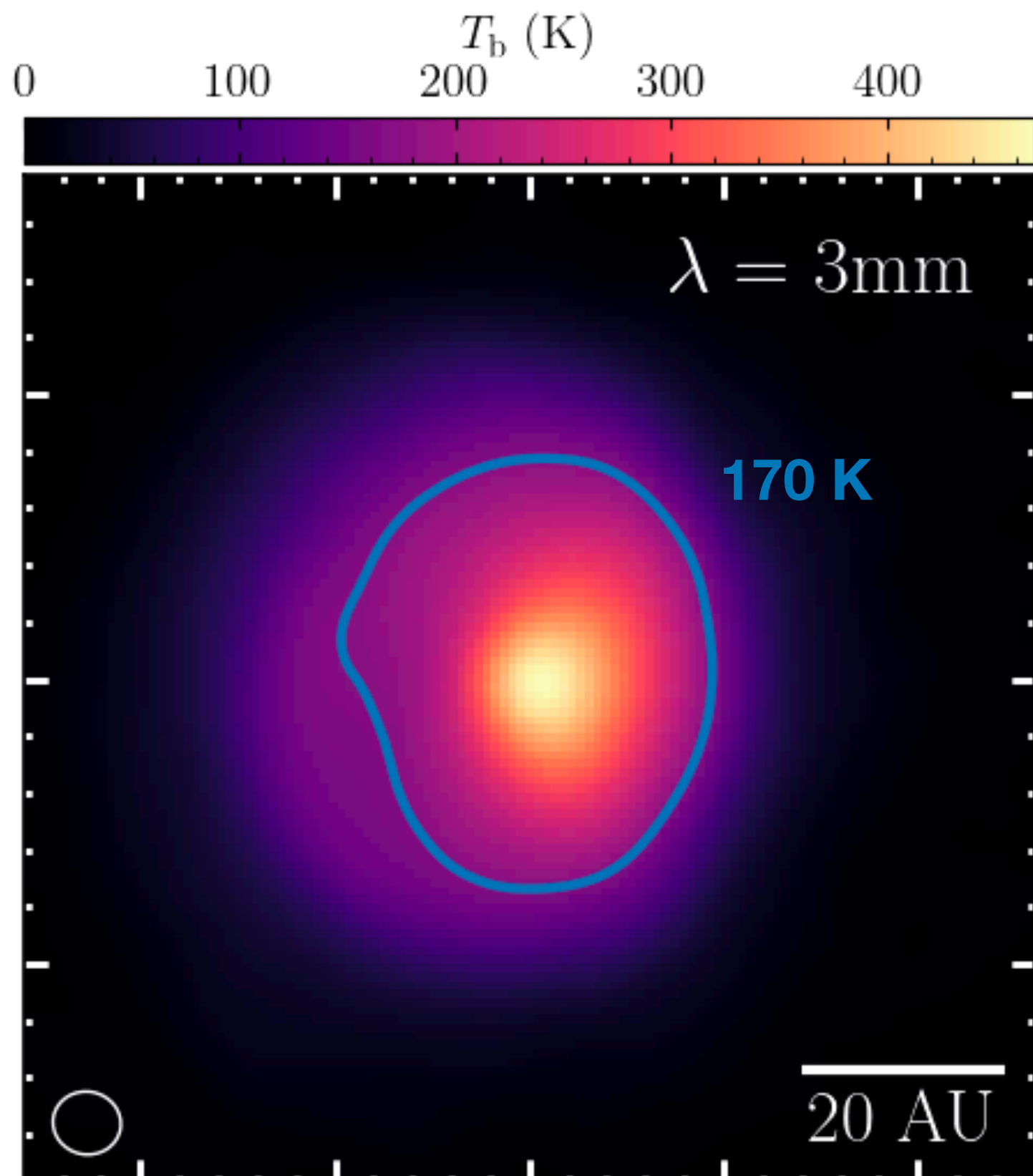
Jorgensen et al. 2016, Maureira et al. 2020



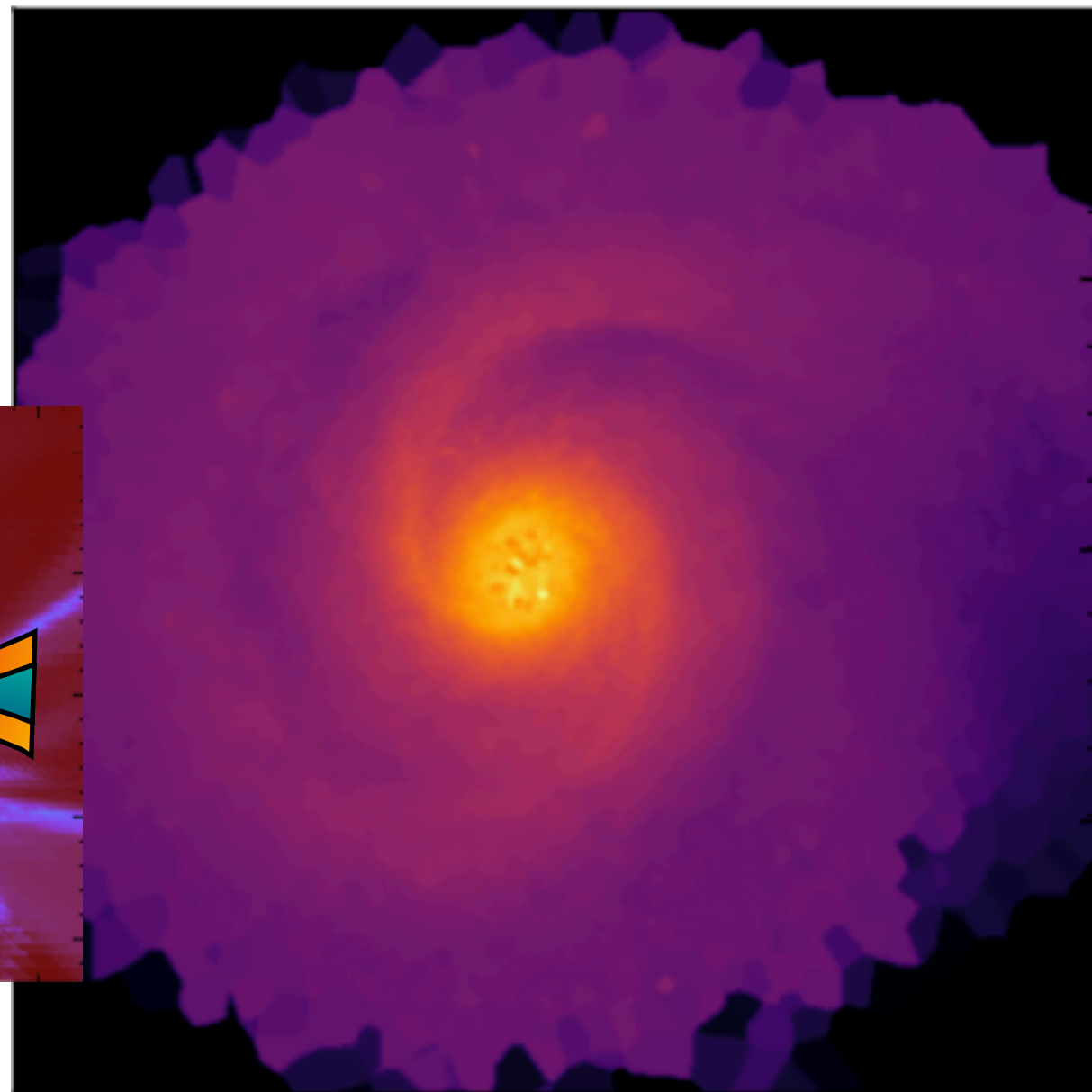


Zamponi, Maureira et al. 2021





Gas temperature (RHD model)



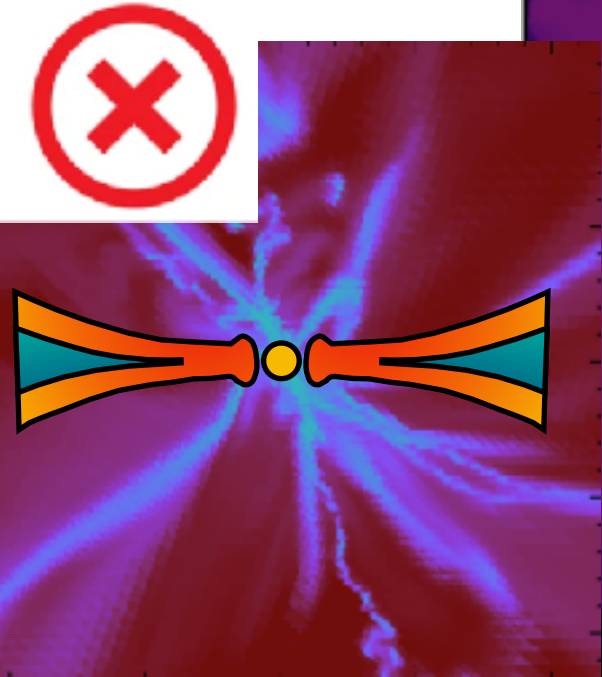
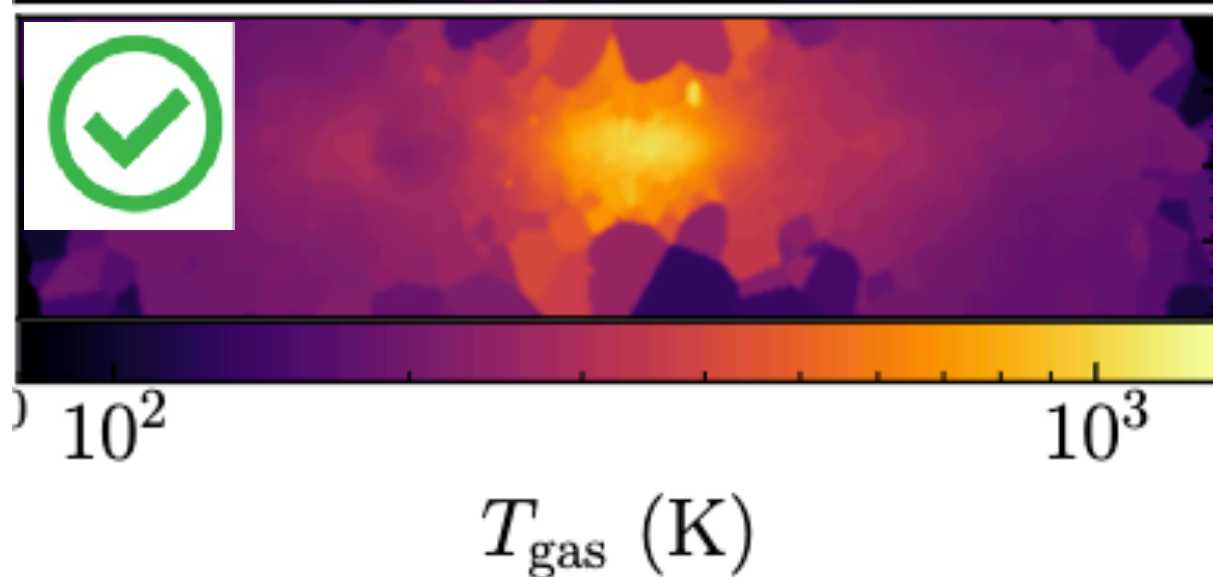
← Reproduces fluxes and produces asymmetries

Gas temperatures needed to match observations

Protostellar radiation Irrelevant (disk too optically thick)

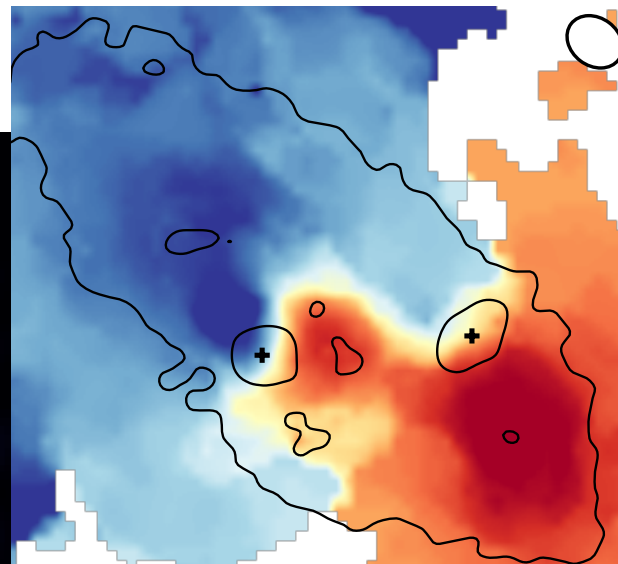
Spec index naturally explained by reverse temp gradient and high optical depth

$a_{\text{max}} \sim 10\mu\text{m}$ - few $100\mu\text{m}$



ALMA 3 mm at 6.5 au resolution

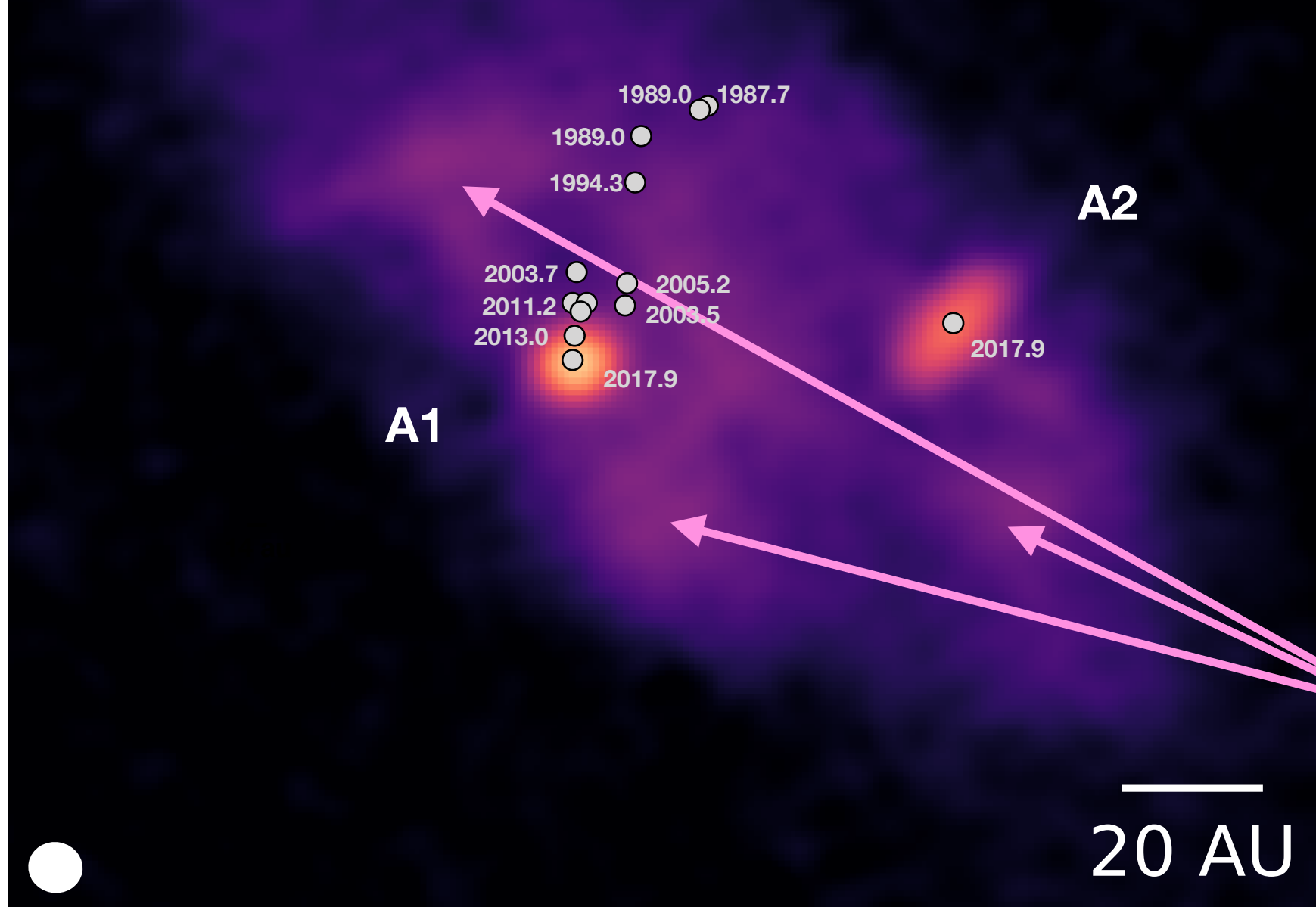
Gas kinematics



Bound system

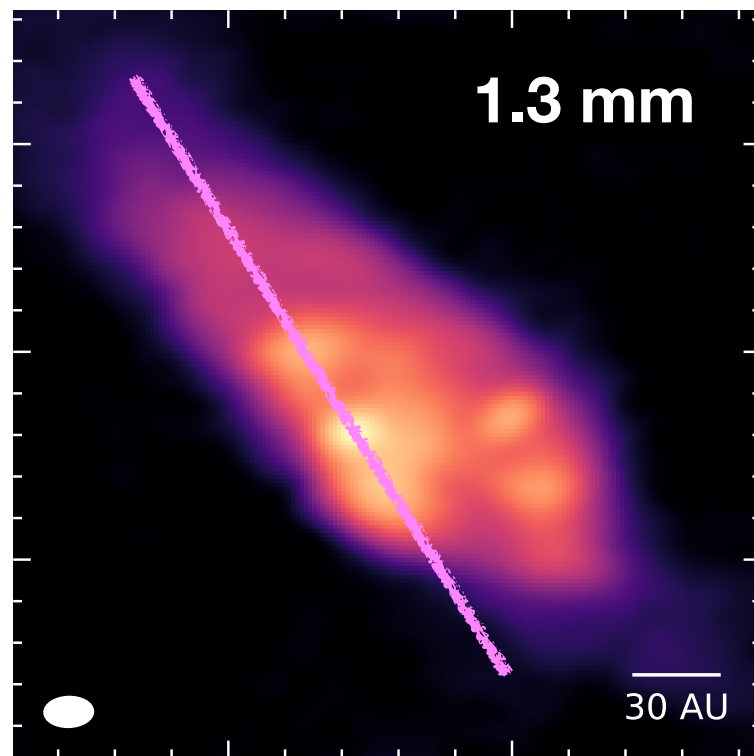
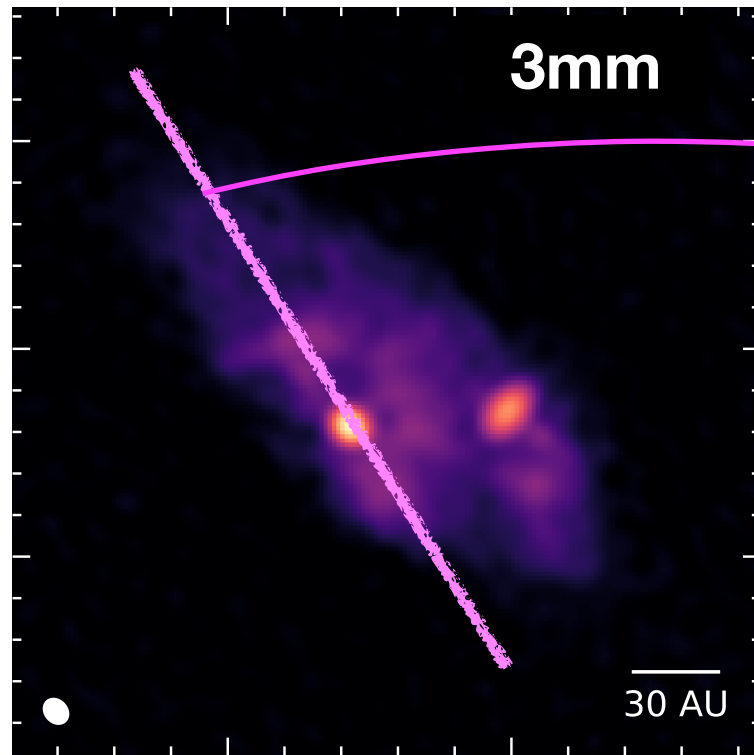
At least 2 Msun in protostellar mass

Stellar kinematics

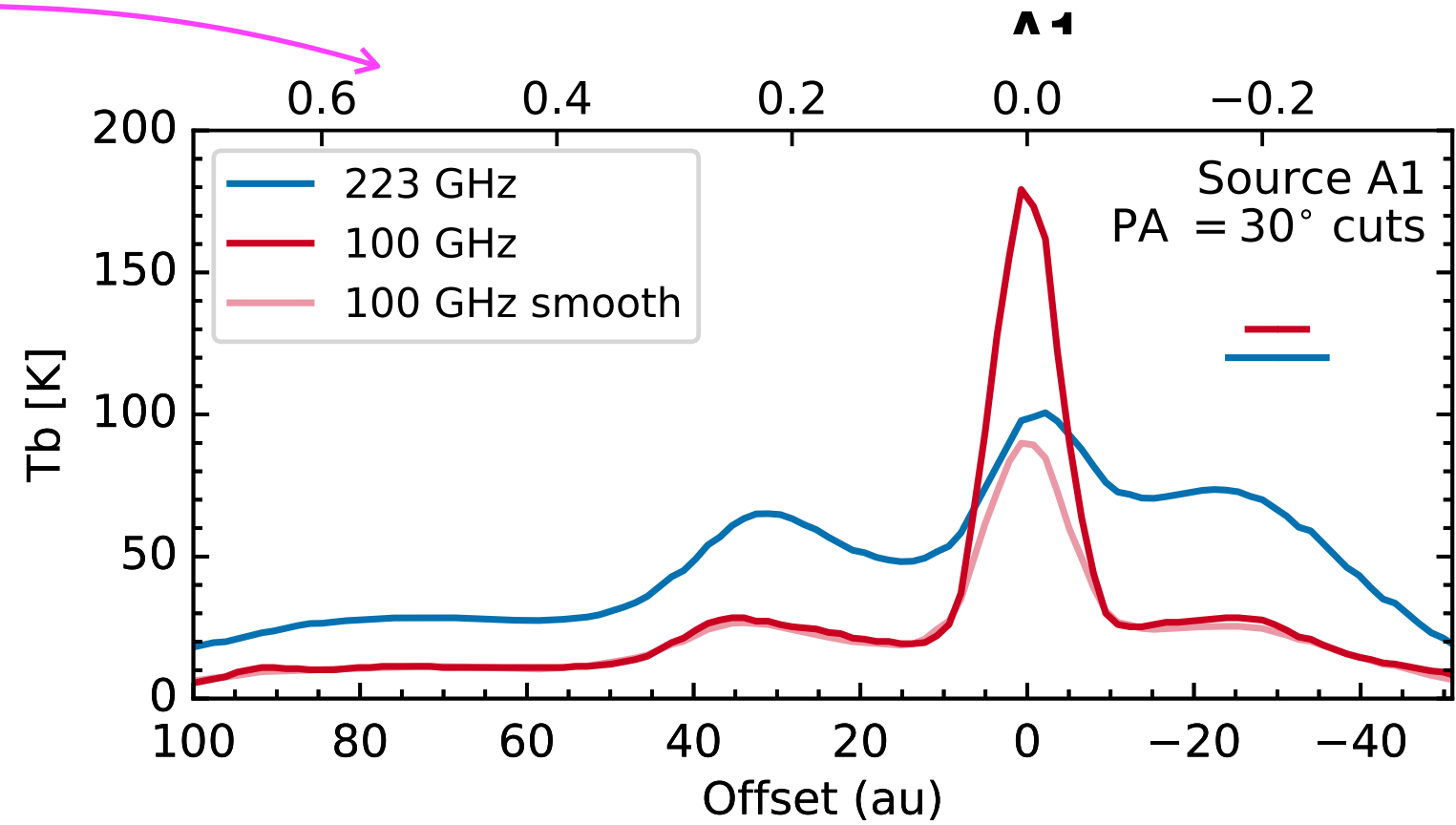


Complex circumbinary substructures

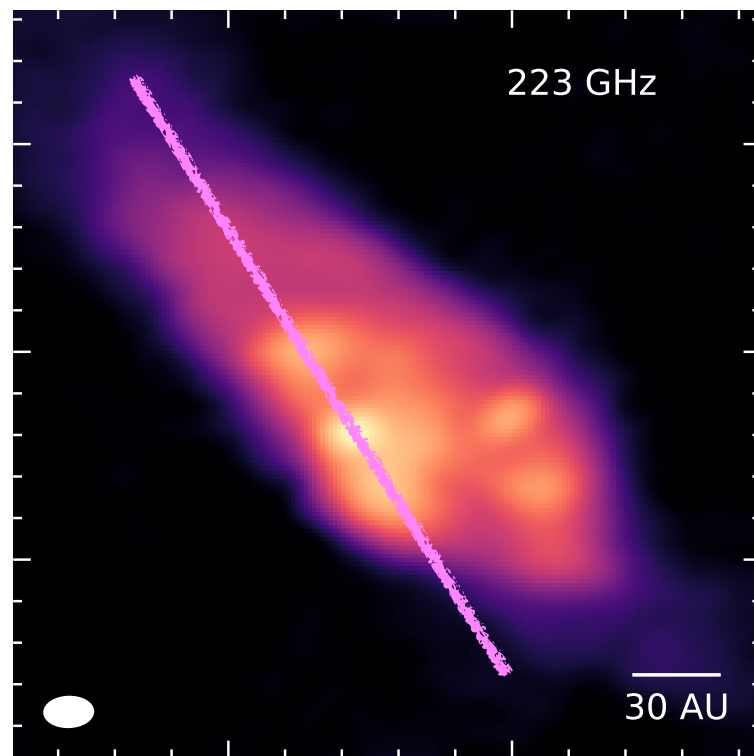
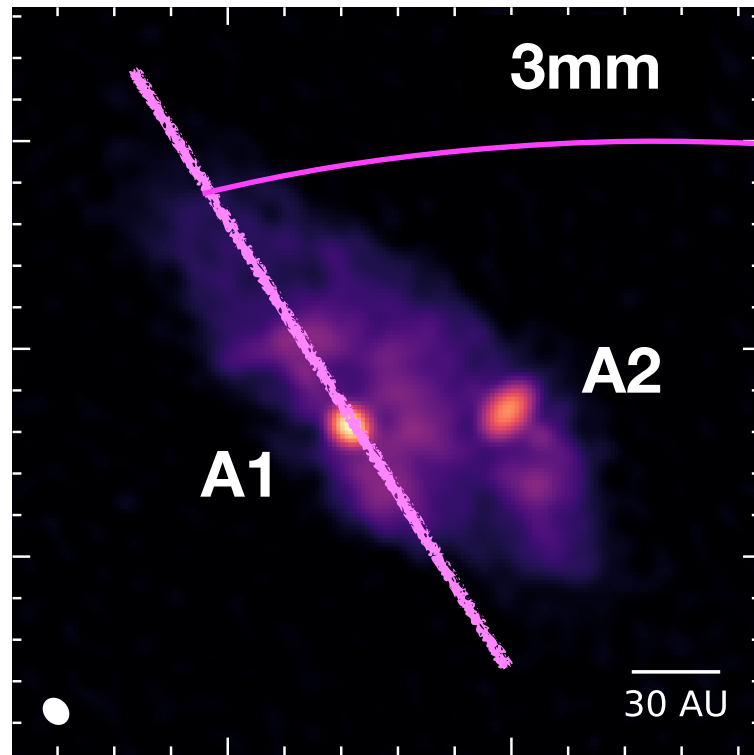
Comparison between spectral index and Brightness Temperature distribution



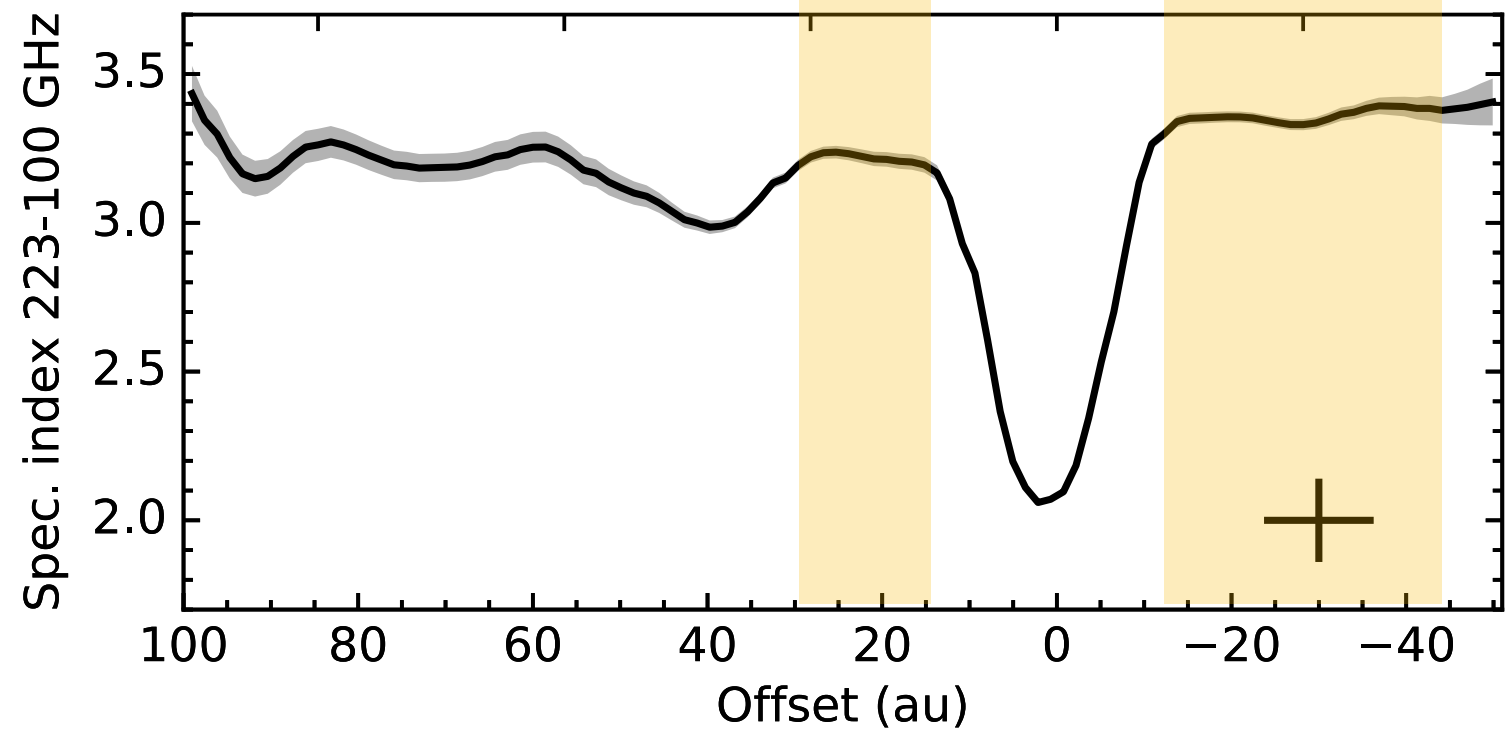
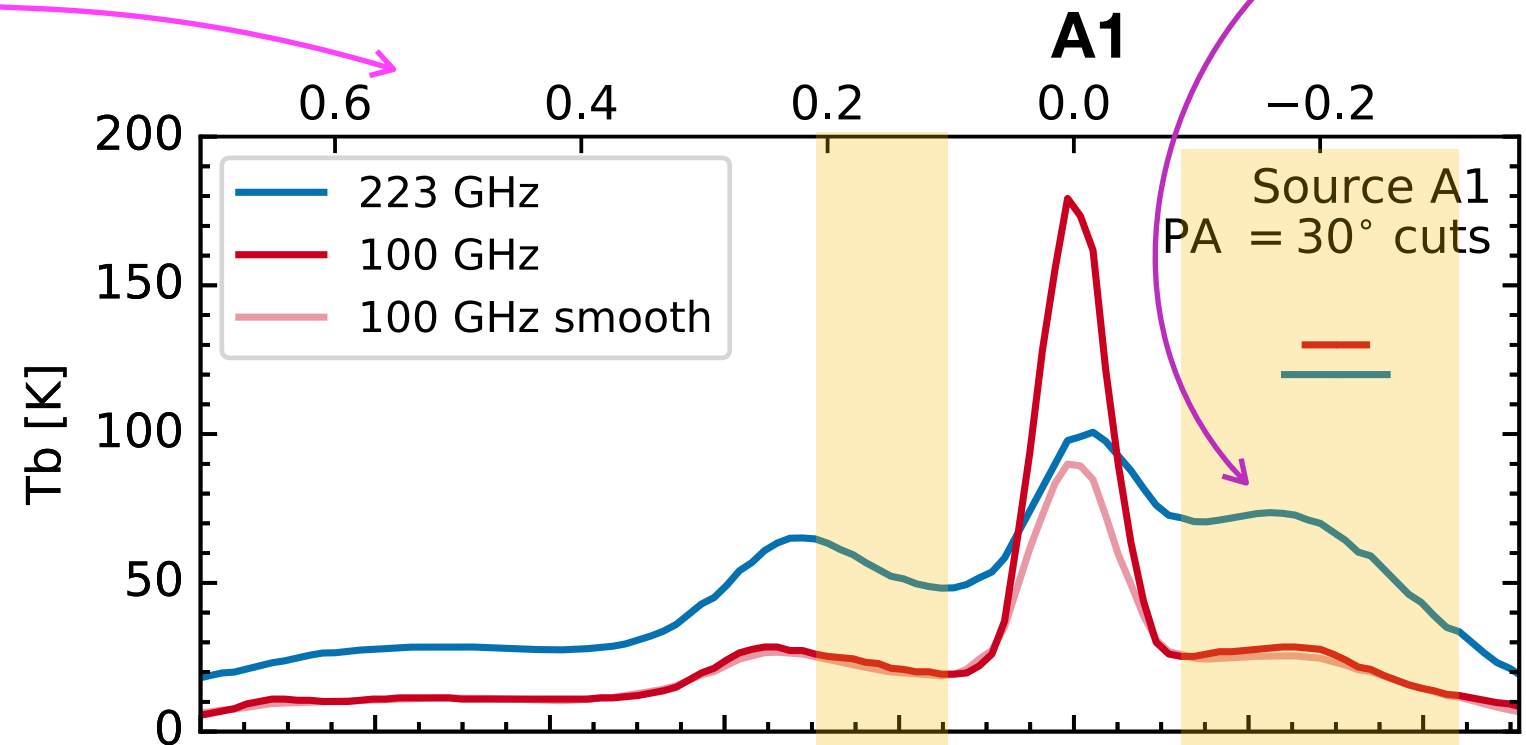
Profiles



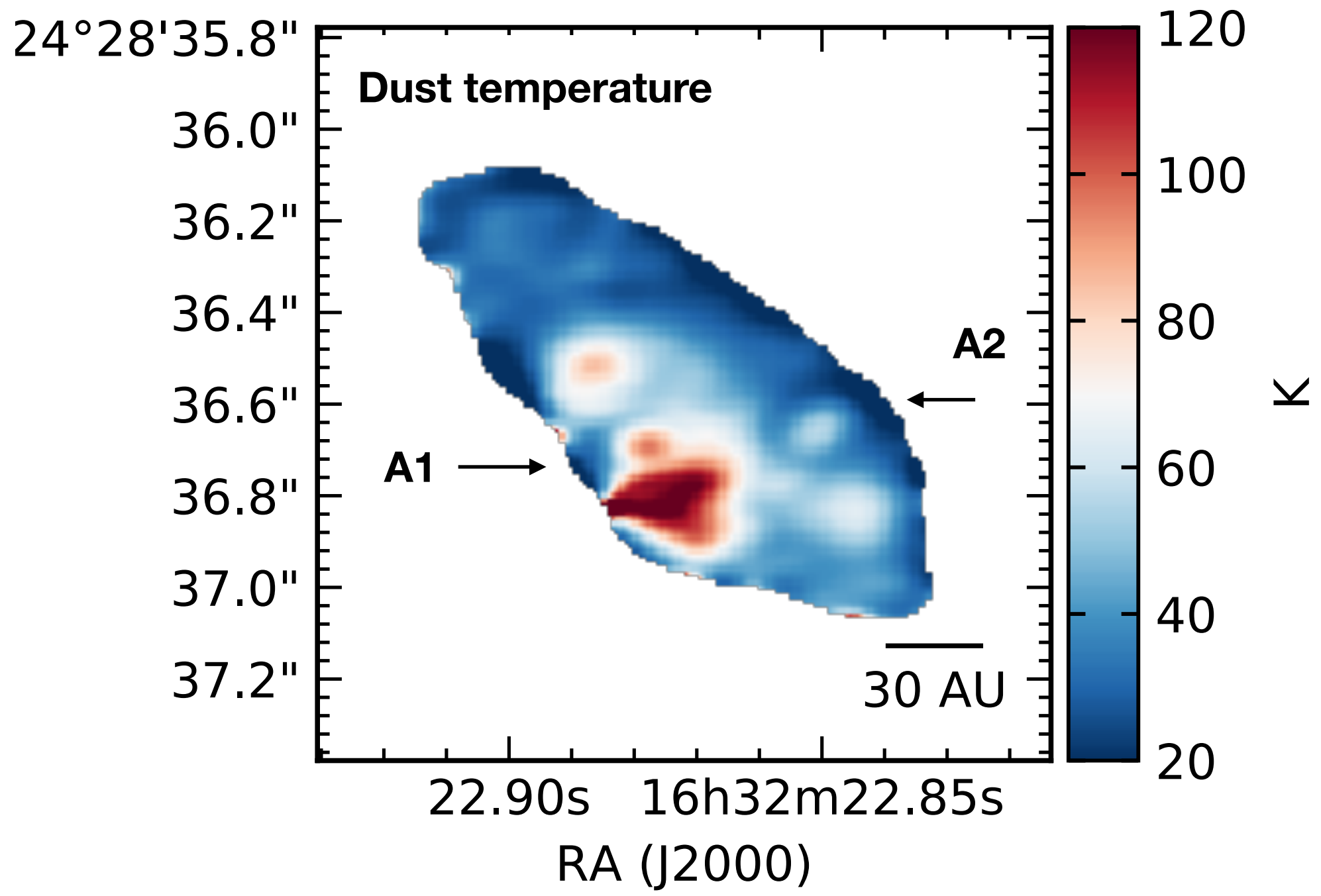
Comparison between spectral index and Brightness Temperature distribution



Profiles

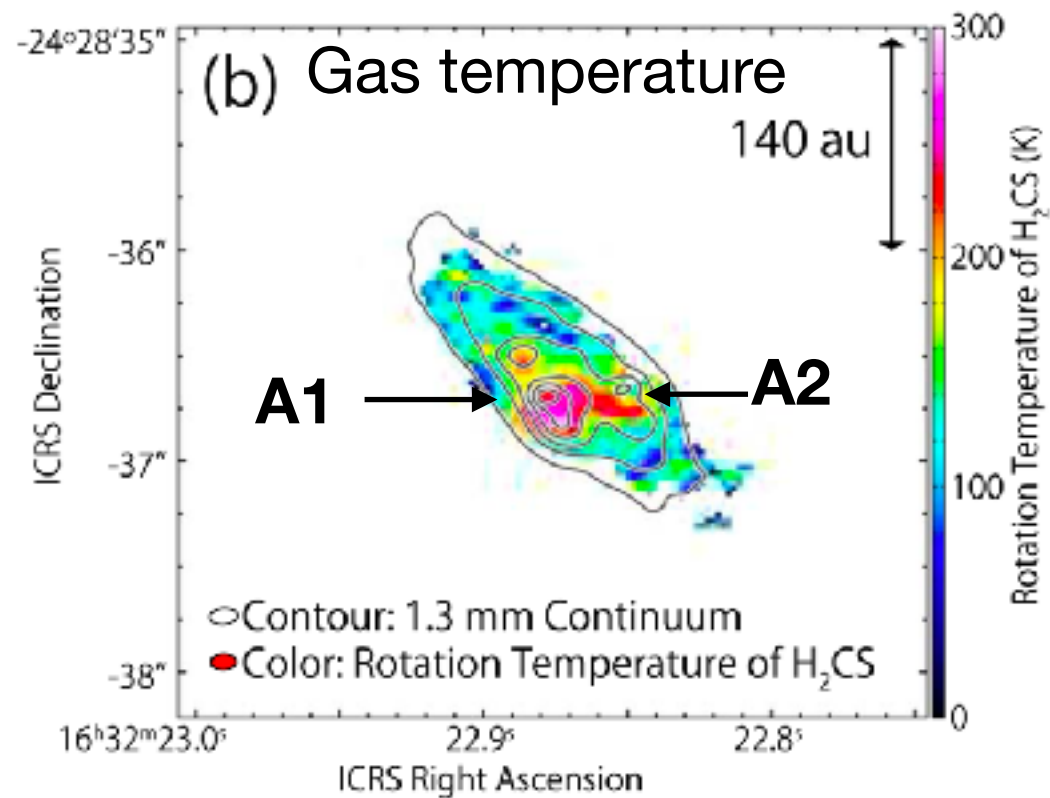


Dust temperature distribution NOT a smoothly decreasing with distance to the protostars

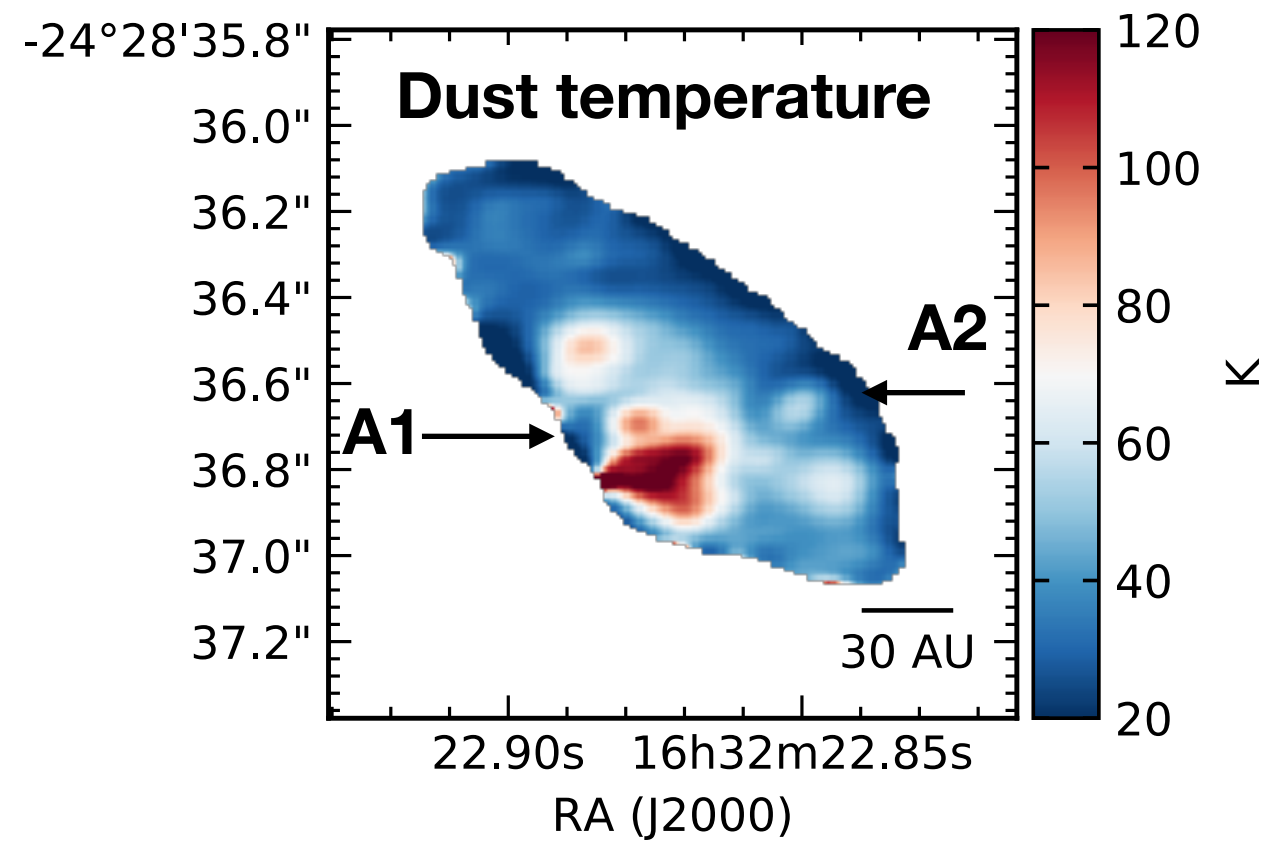


*Lower limits assuming beta=1.7

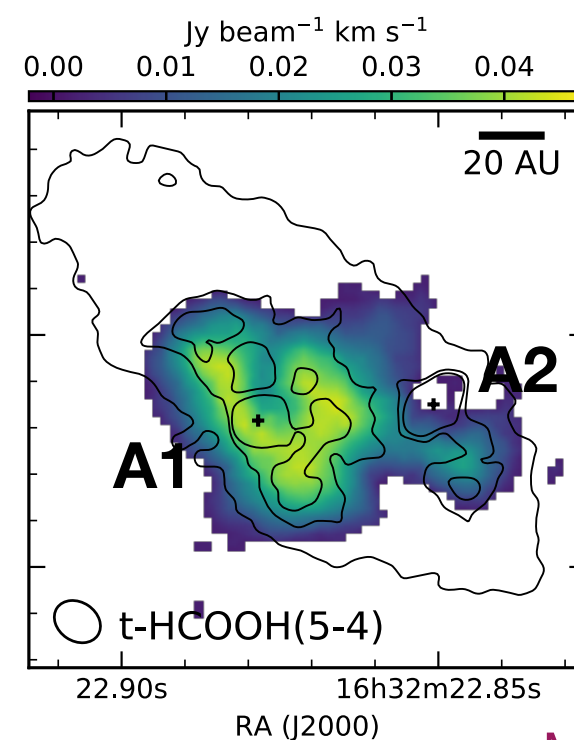
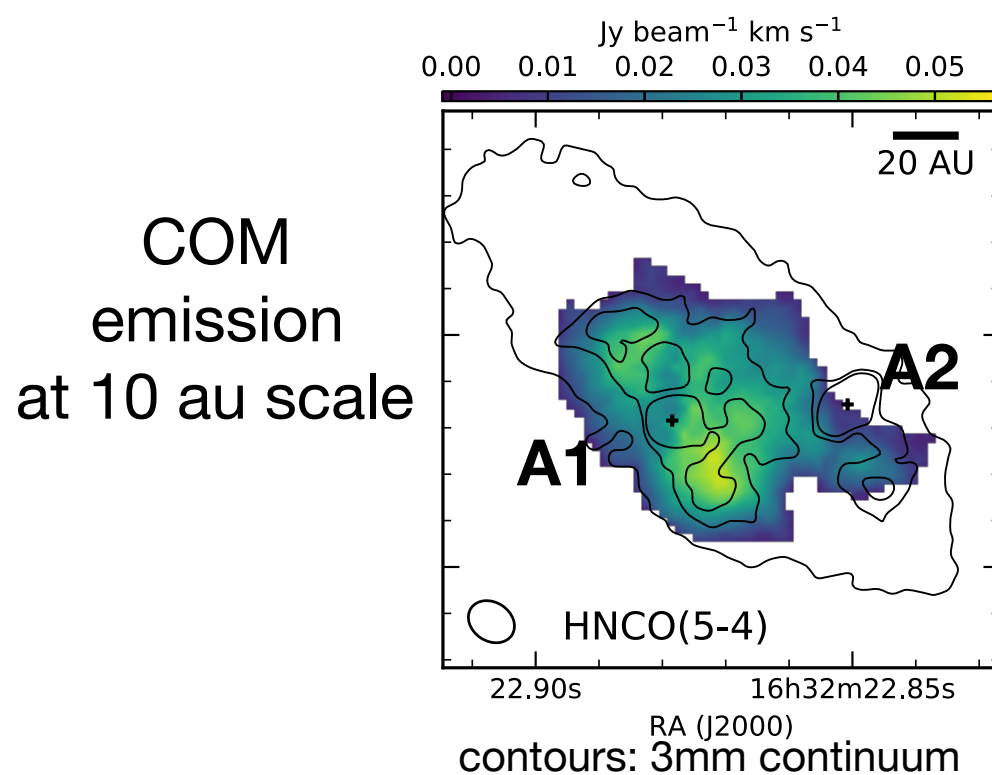
Gas and dust show localized enhancements of temperature



Oya et al. 2020 (see also van't hoff et al. 2020)



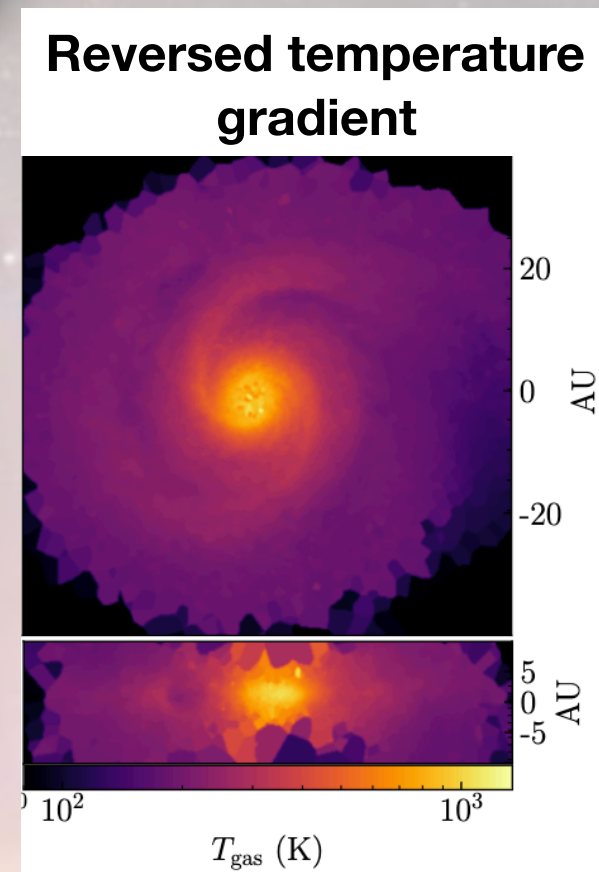
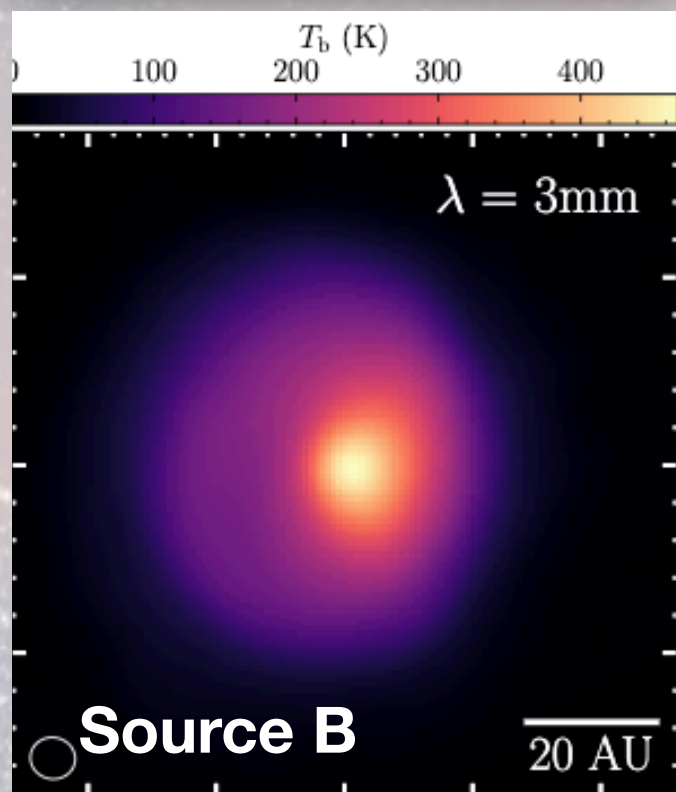
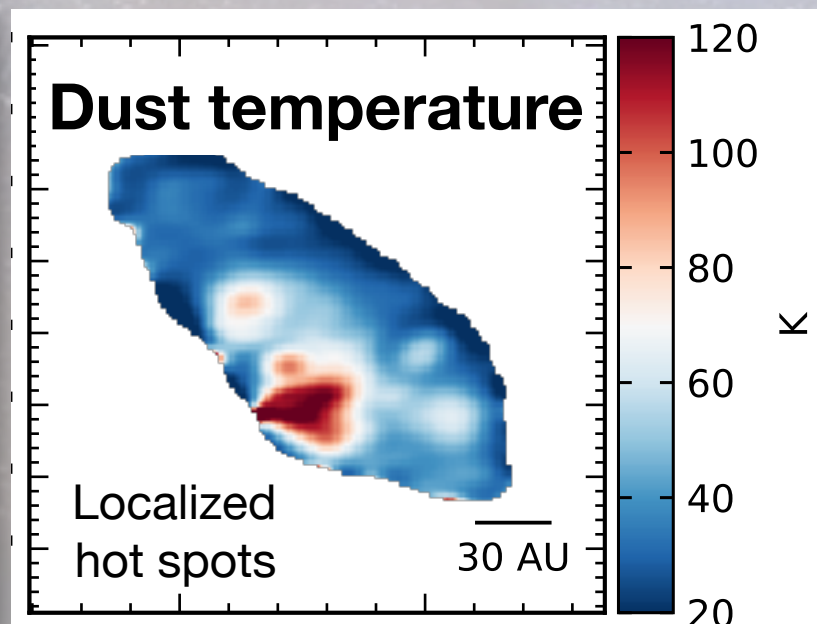
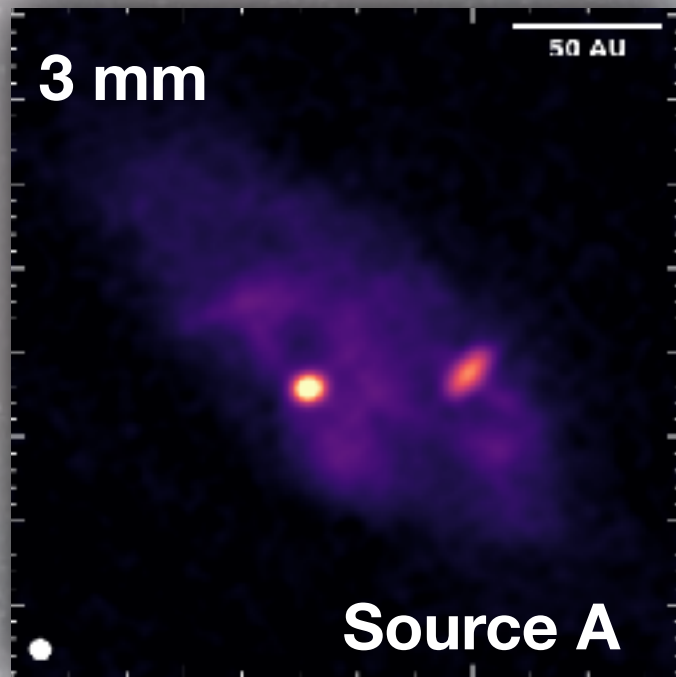
Maureira et al. (to be submitted)



Protostellar radiation cannot explain the temperature distribution in the gas and dust (shocks???)

Maureira et al. 2020

Conclusions and questions



Passively radiated disk
not good representation of the
disk structures
in the Class 0 IRAS 16293-2422

What about others? More obs
please! (Lin et al. 2020 HH212)

Implications for calculations of
masses, and interpretation of
spectral index (Li et al. 2017, Galvan-
Madrid et al. 2018)

What are the implication for
planet formation?

